

Retrieval of thermospheric density and wind from space-borne accelerometers

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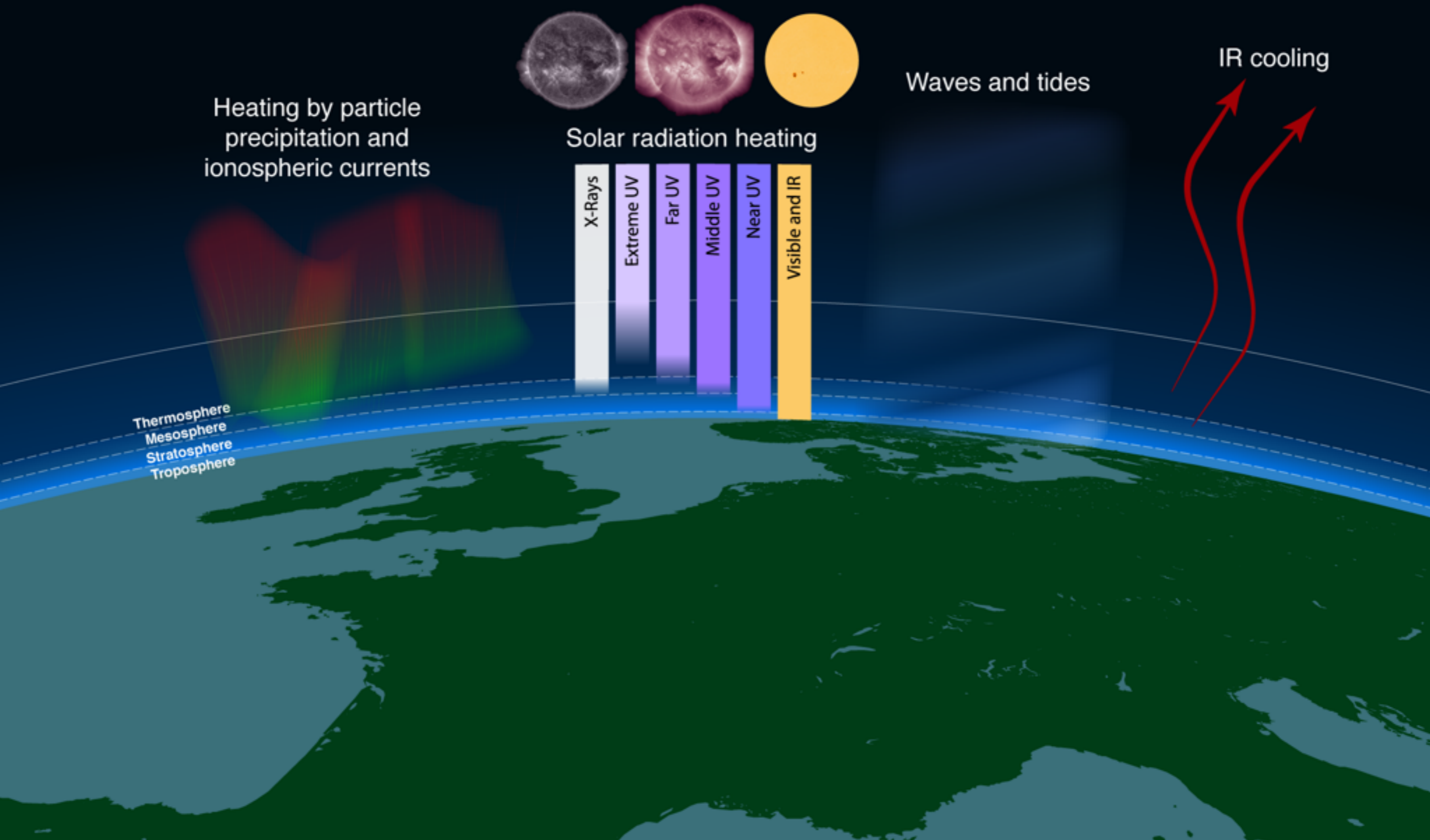
Contents

- Introduction to the thermosphere
- Experience with CHAMP, GRACE, GOCE and Swarm data processing
- Outlook and requirements for MICROSCOPE

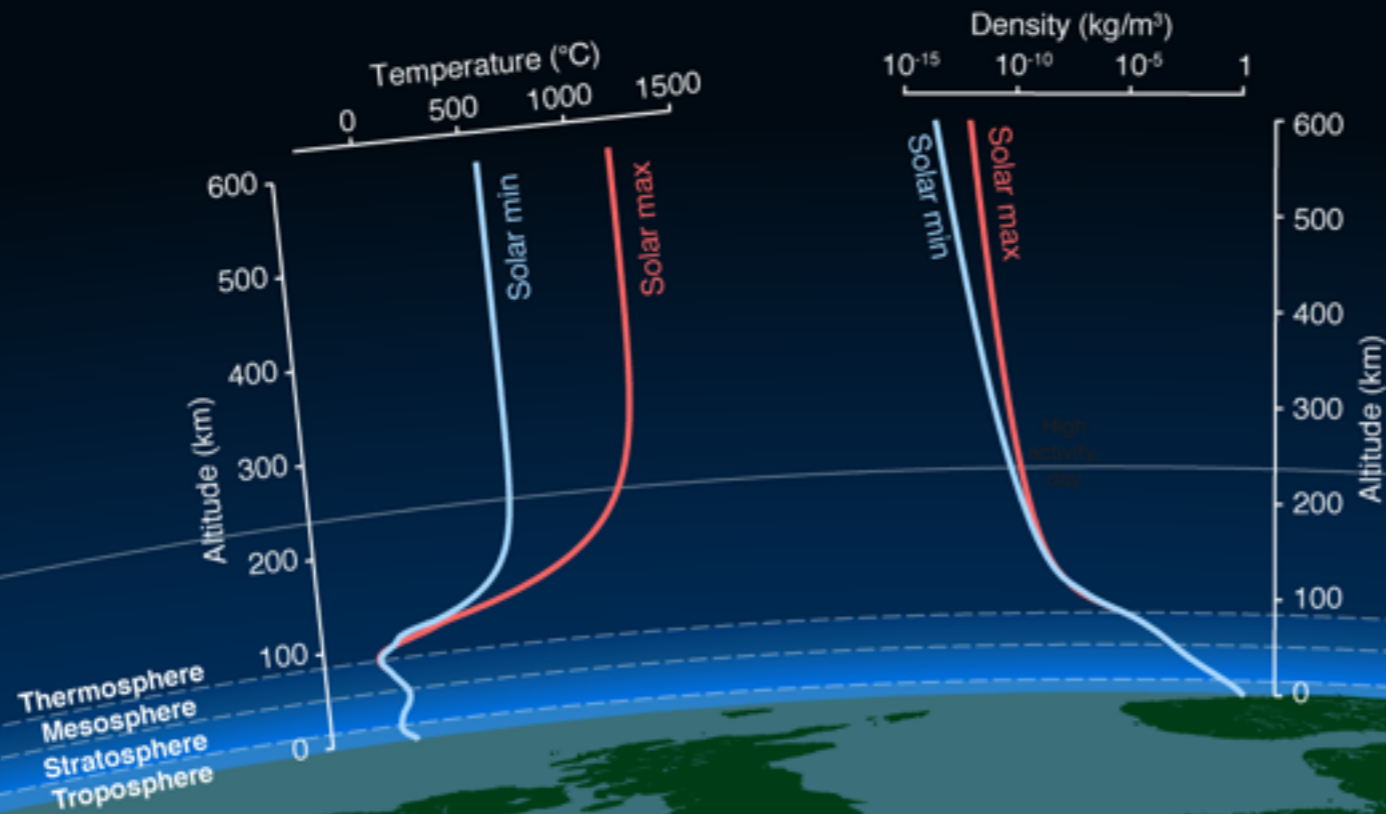
The thermosphere and exosphere

- Upper layers of the Earth's atmosphere, starting at ~90 km.
- Strong coupling of thermosphere dynamics with ionosphere and magnetosphere.
- Large thermospheric variability, driven by solar UV radiation, Joule heating and particle precipitation, taking place in the lower thermosphere.

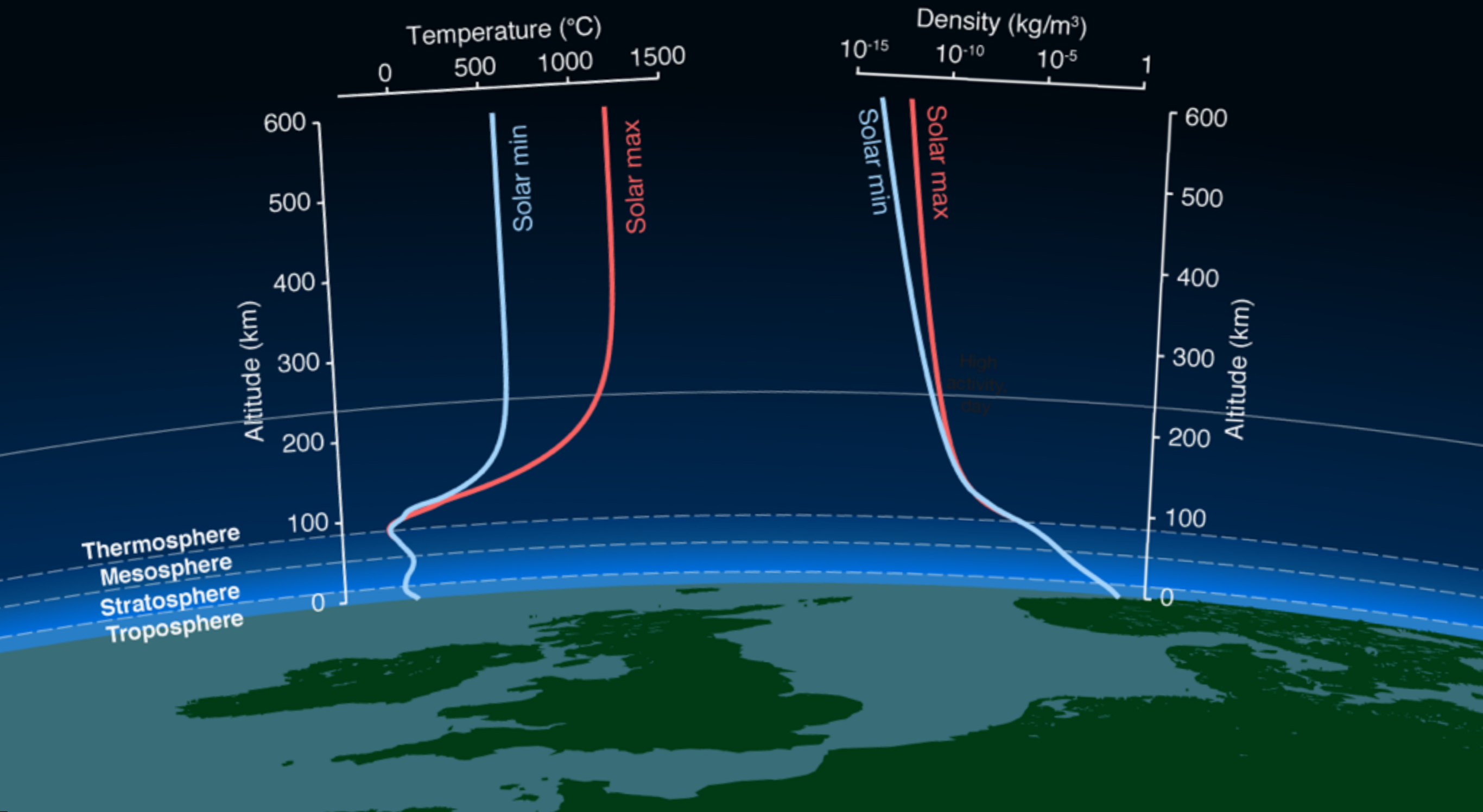
Energy sources

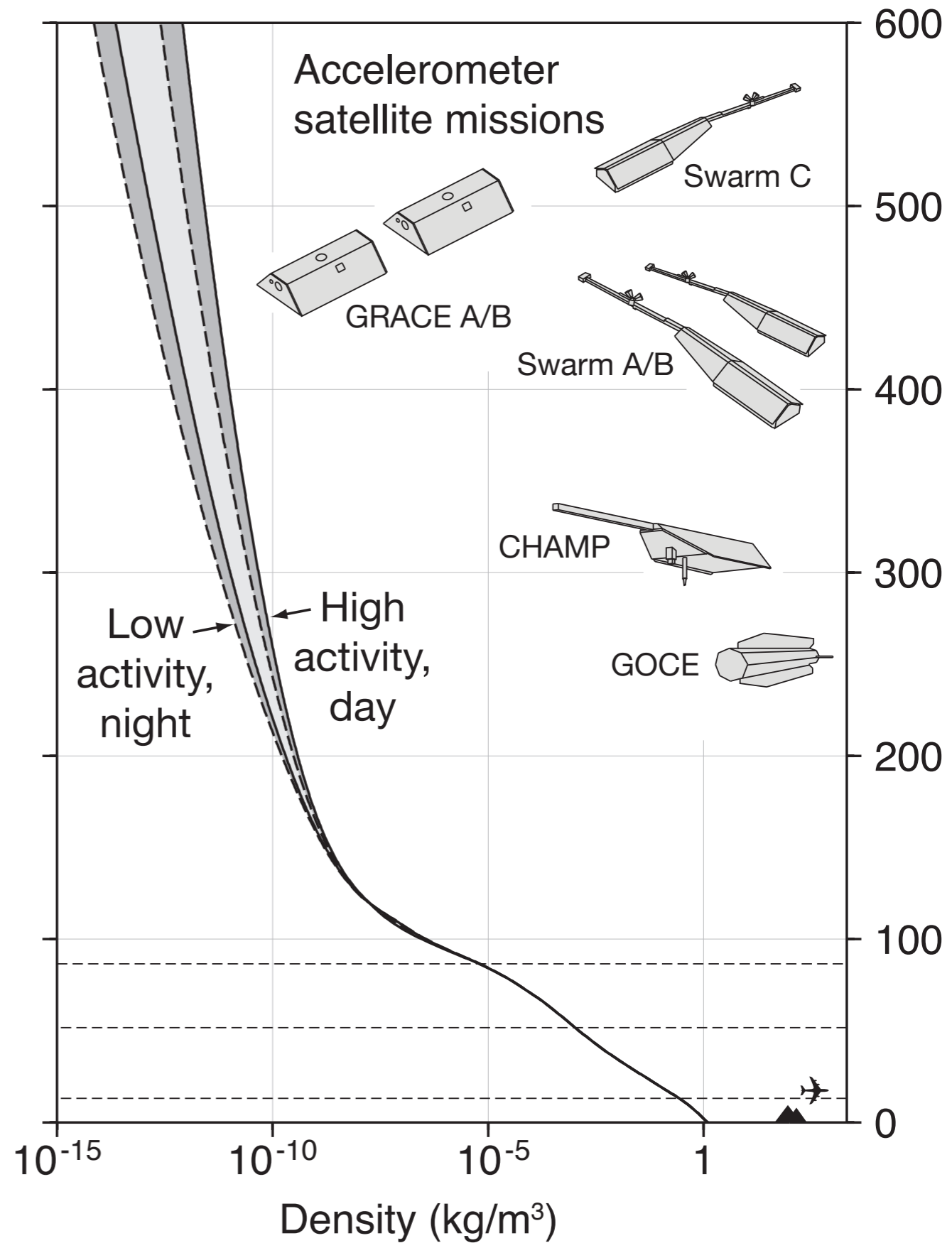
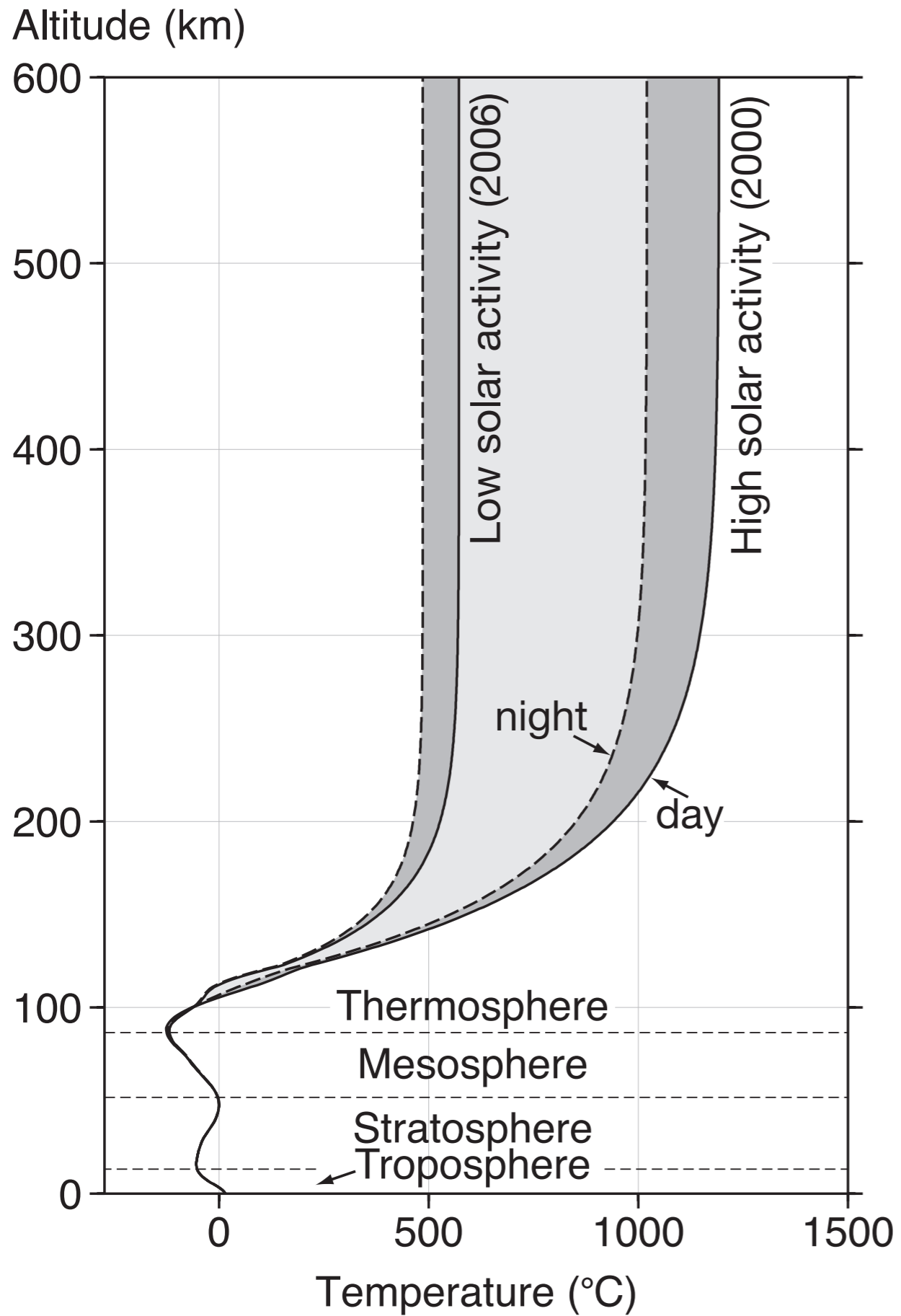


Temperature and density



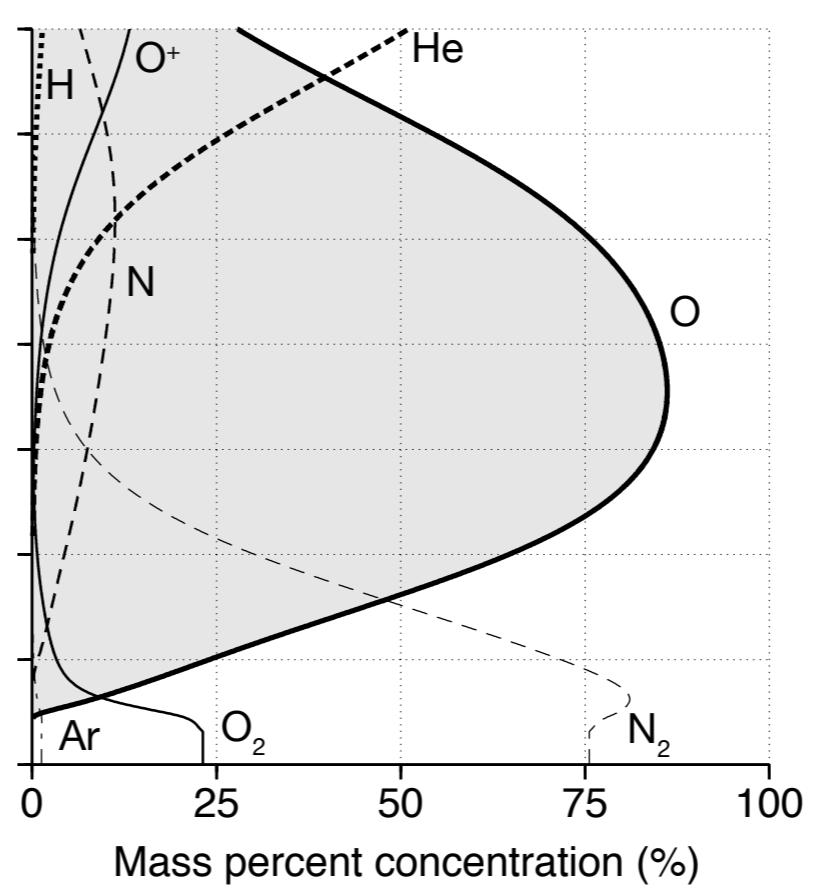
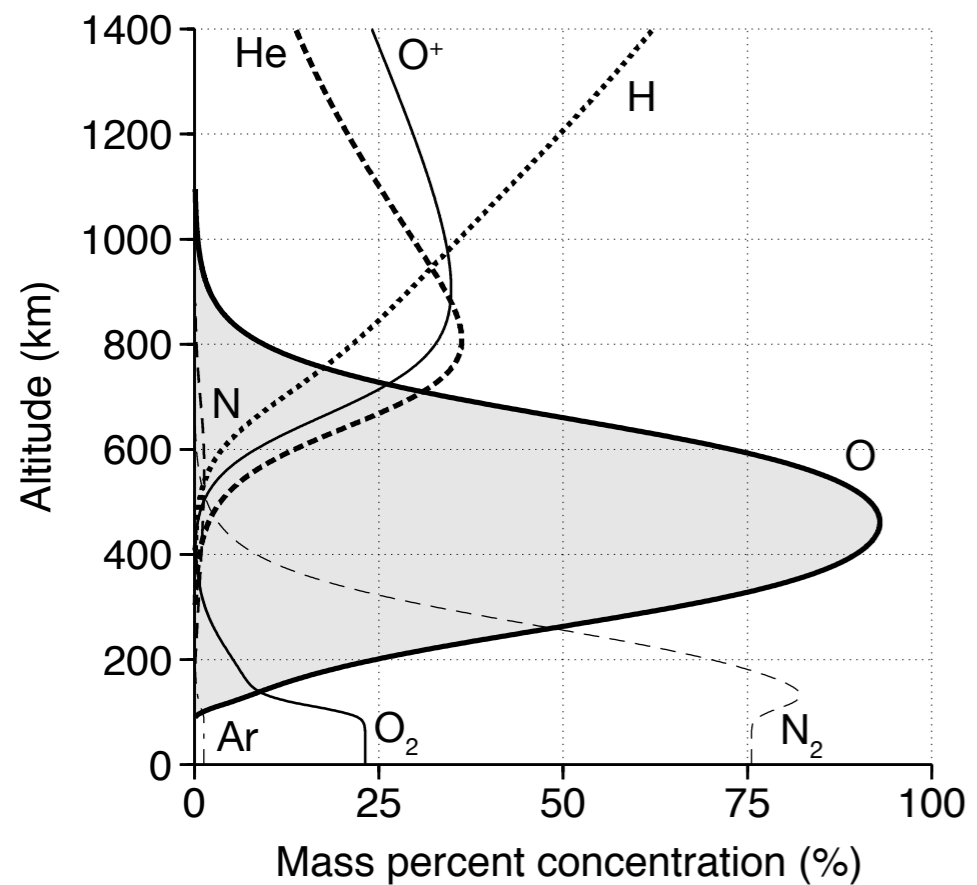
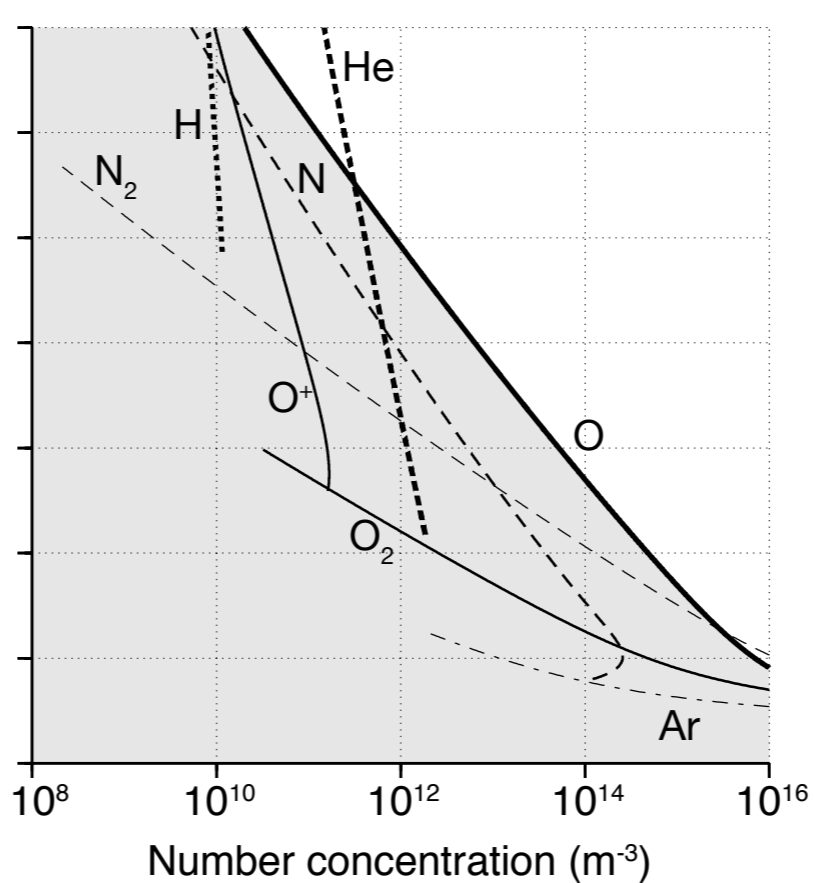
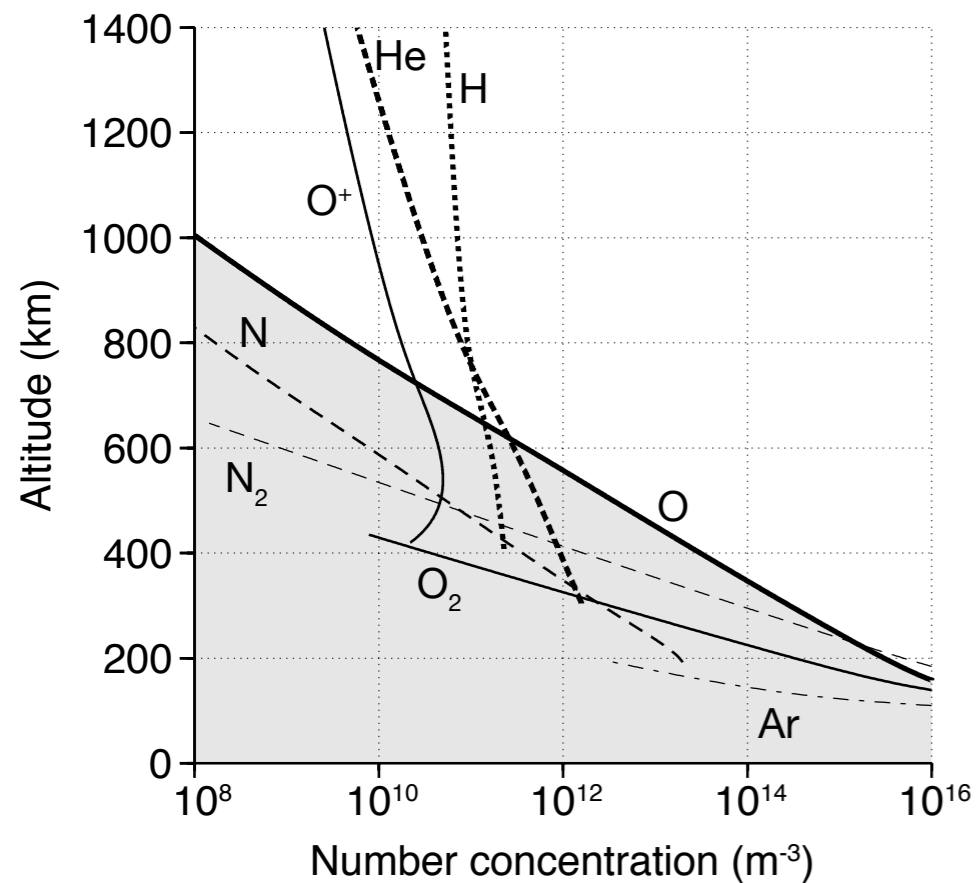
Temperature and density



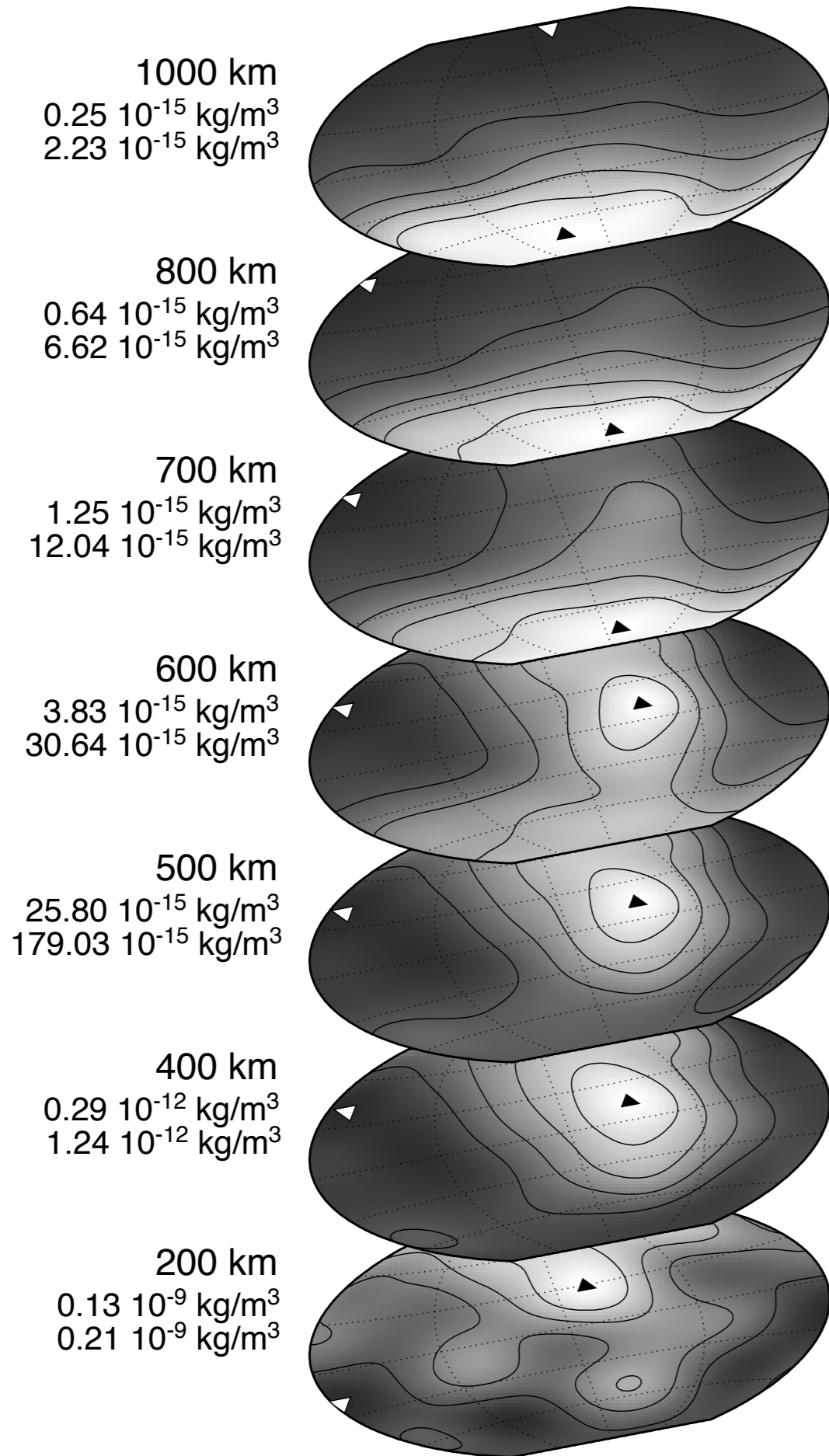


Low activity (2006)

High activity (2000)



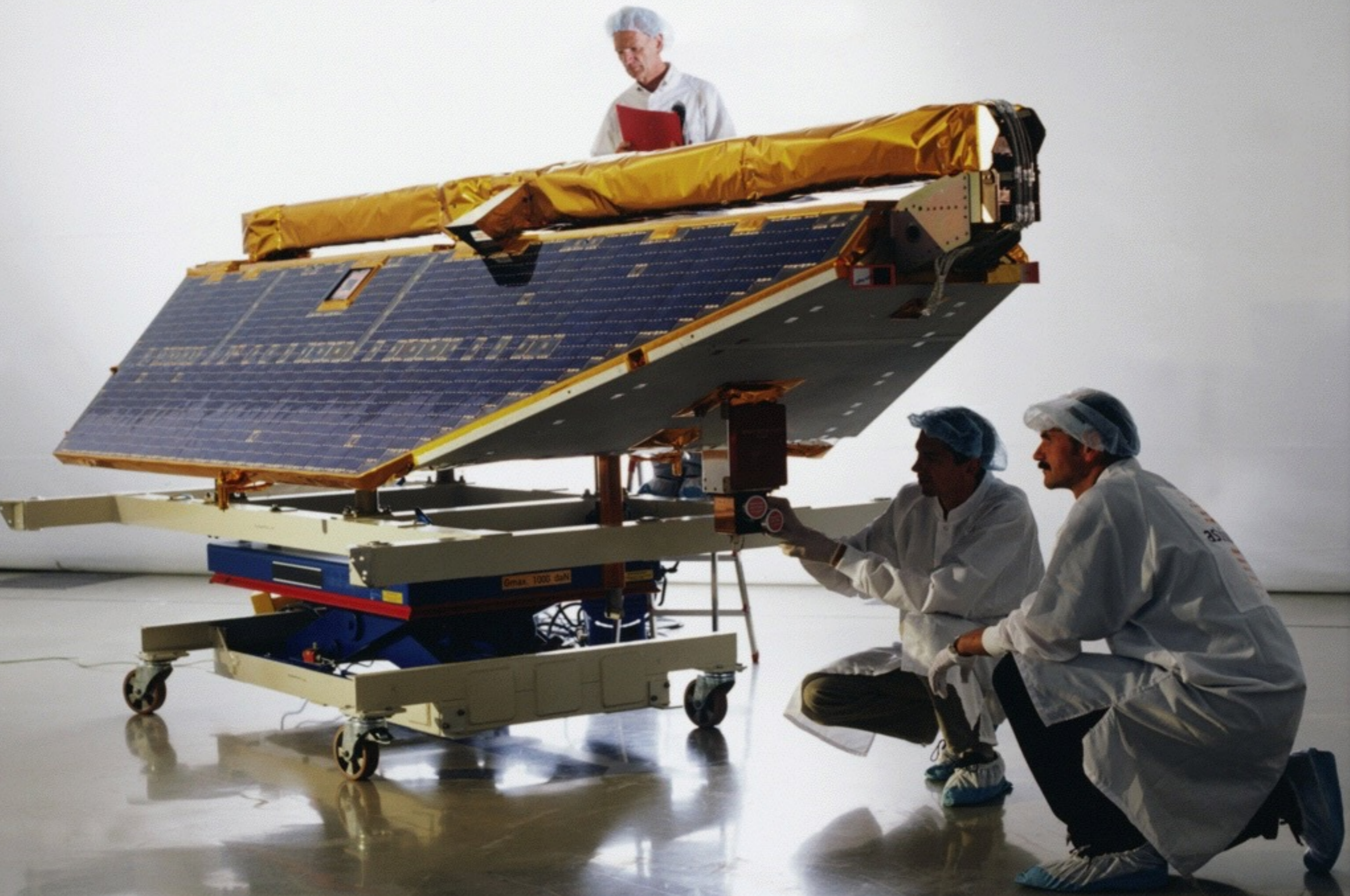
Low activity (2006)



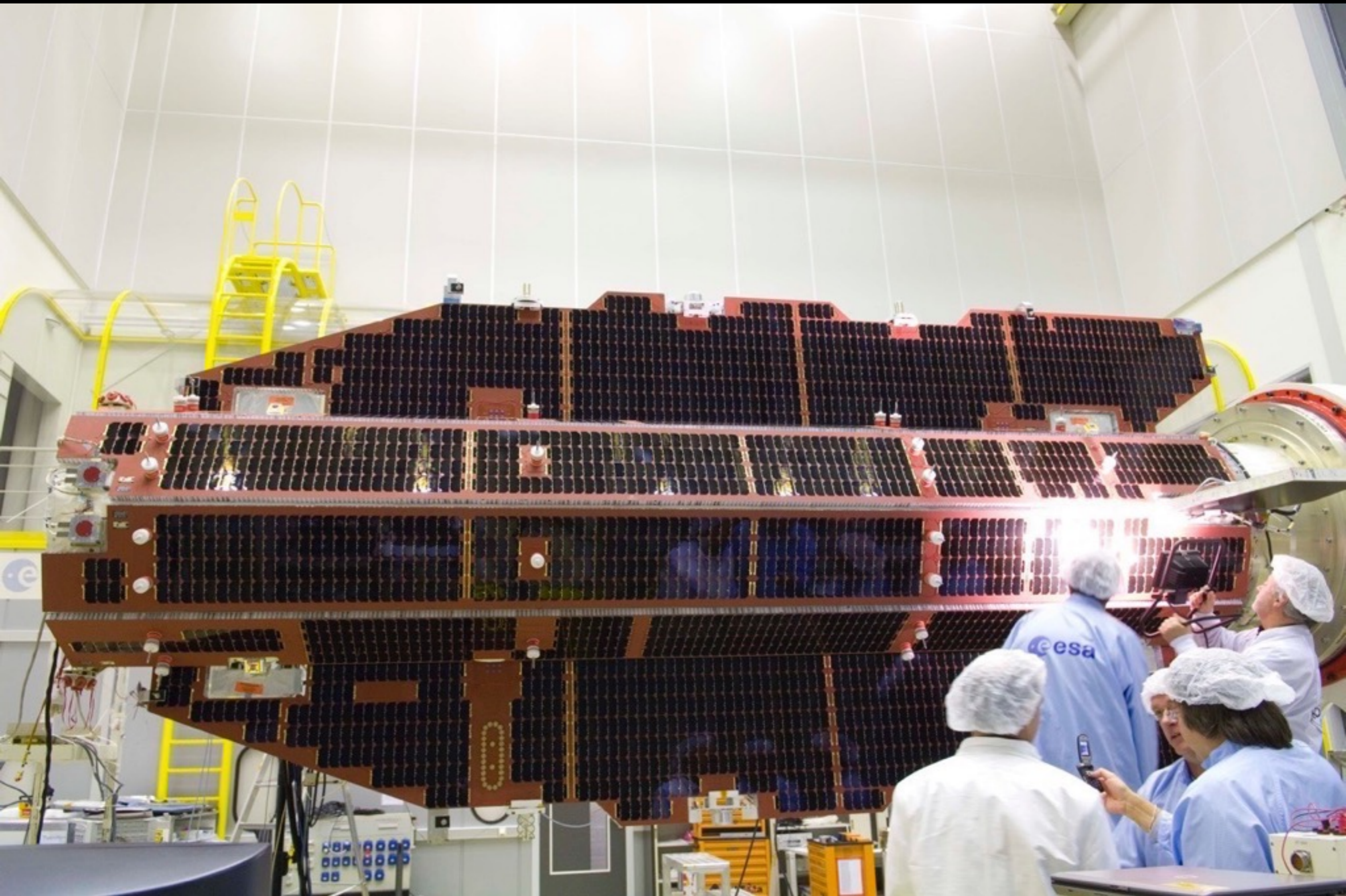
Latitude / local solar time density variation
at low solar activity, according to the
NRLMSISE-00 model.

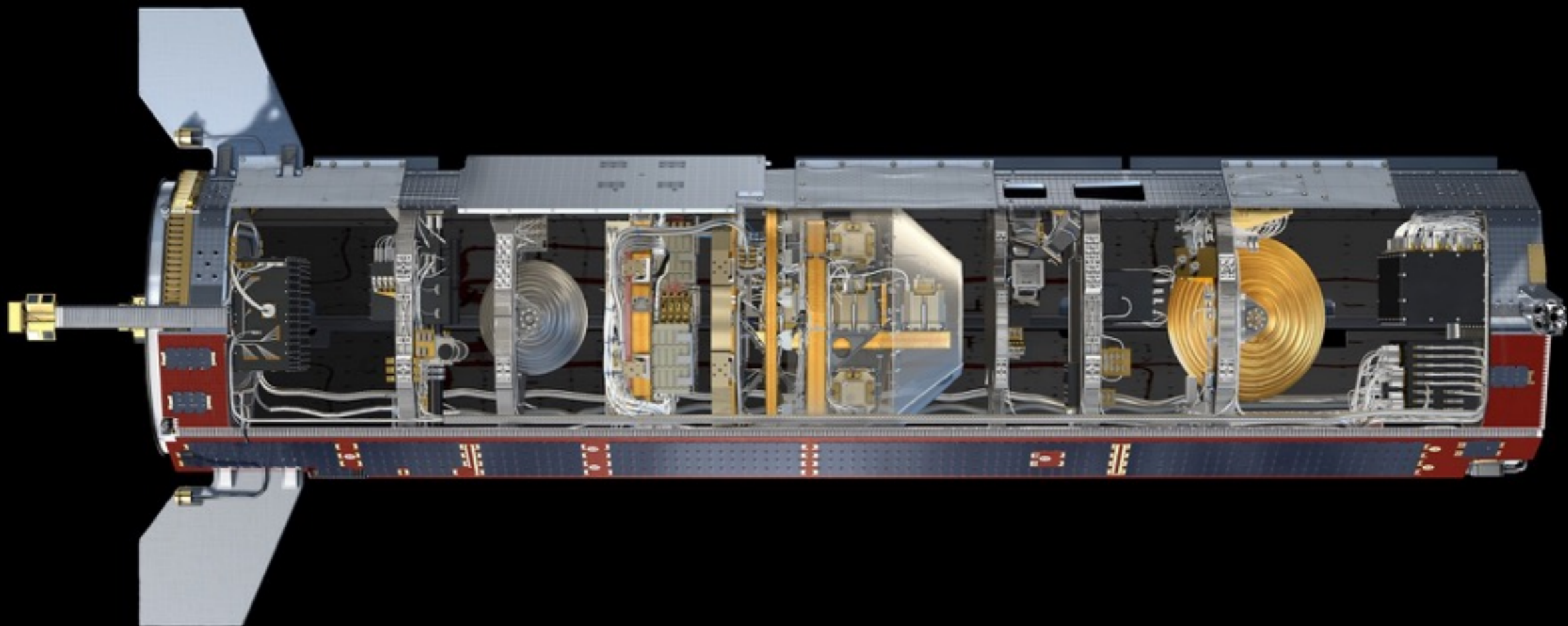
Experience with CHAMP, GRACE, GOCE and Swarm data processing

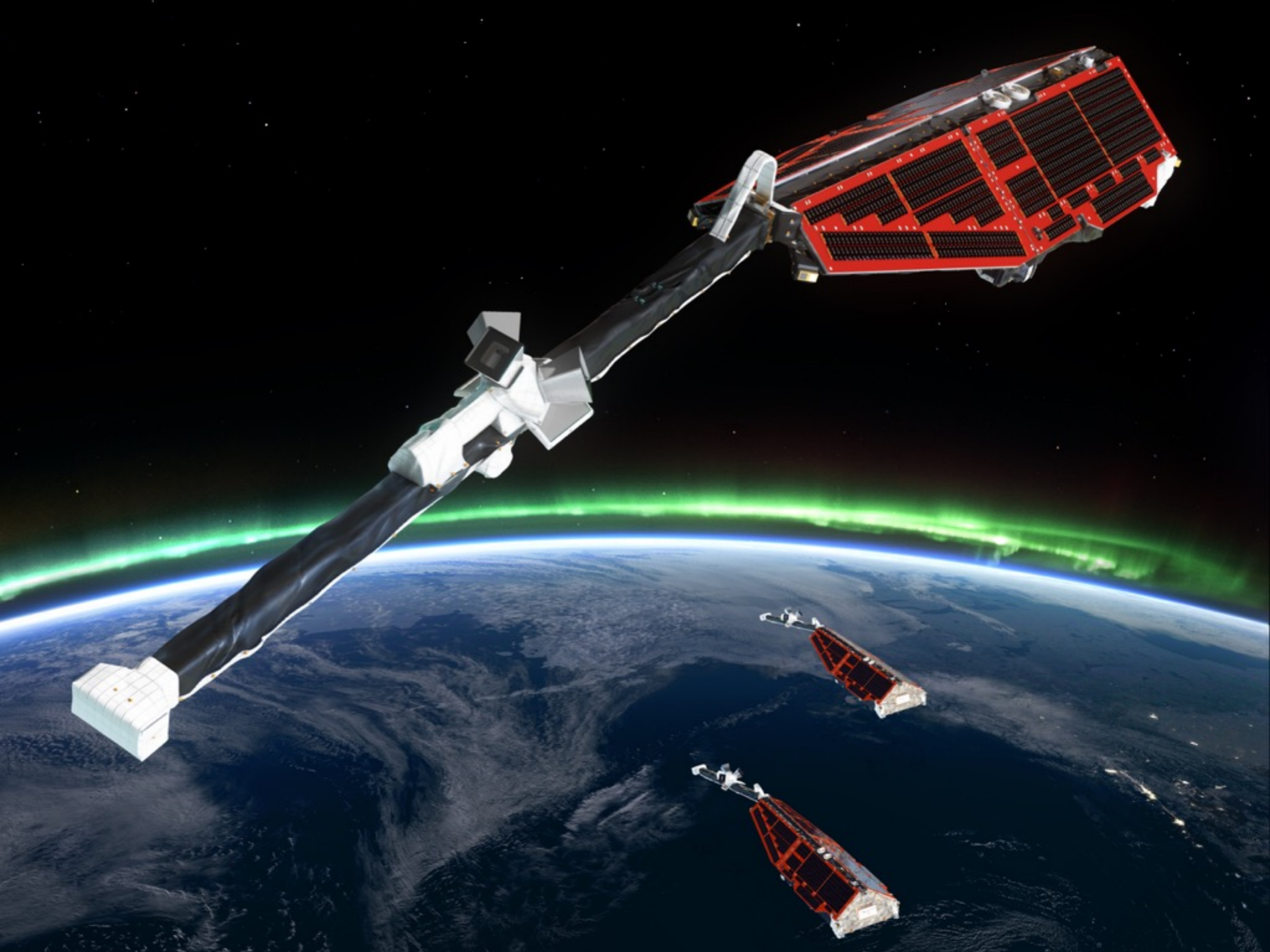
- Goal: create an accurate long-term, high temporal resolution, multi-mission thermosphere data set.
- Application areas:
 - Detailed study of climatology and weather in the thermosphere.
 - Validation of numerical thermosphere-ionosphere models.
 - Input to empirical models for use in space mission planning, operations, re-entry prediction, collision avoidance, etc.

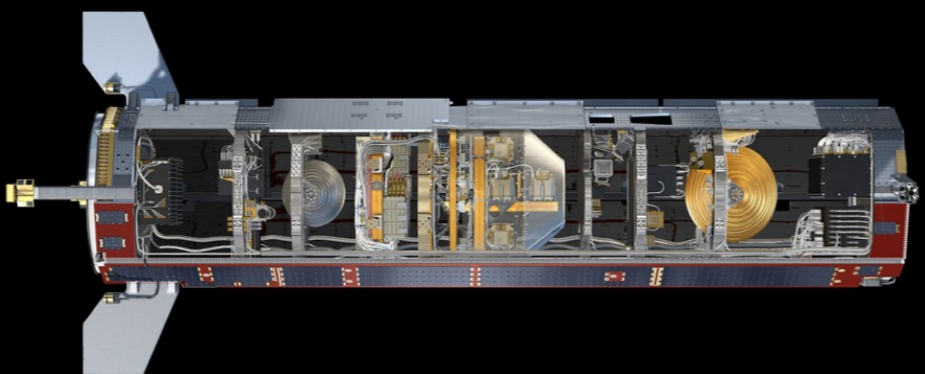








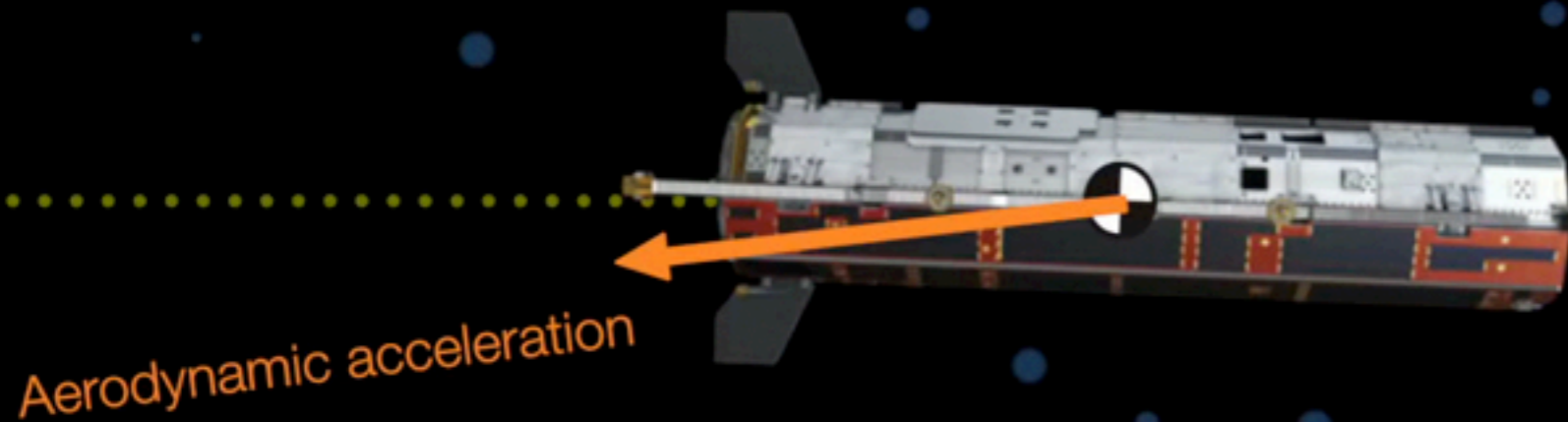




GOCE orbital dynamics: Aerodynamic acceleration

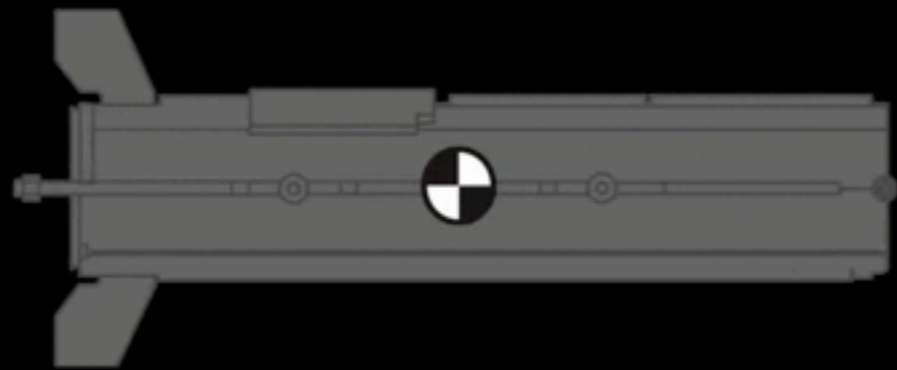


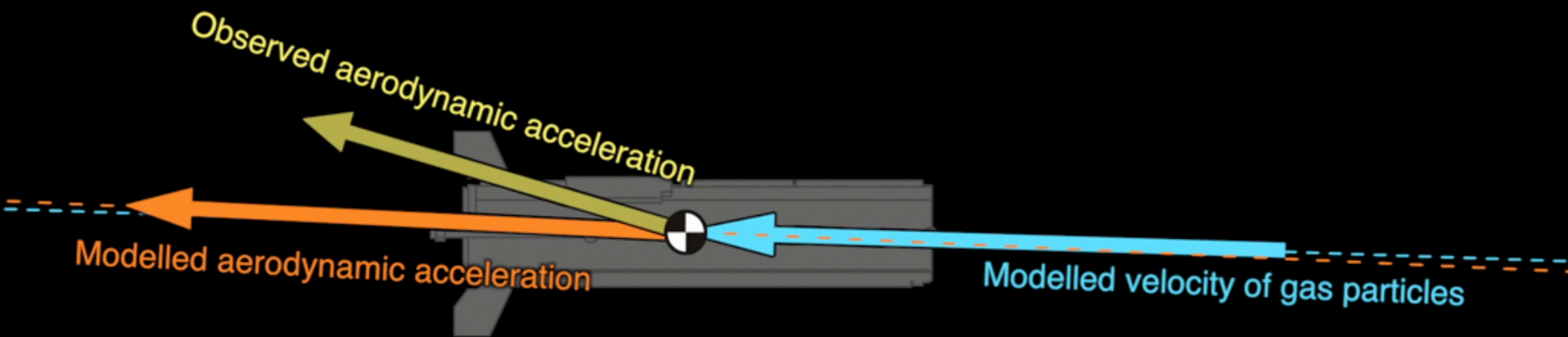
GOCE orbital dynamics: Ion thruster acceleration



GOCE orbital dynamics: Radiation pressure acceleration





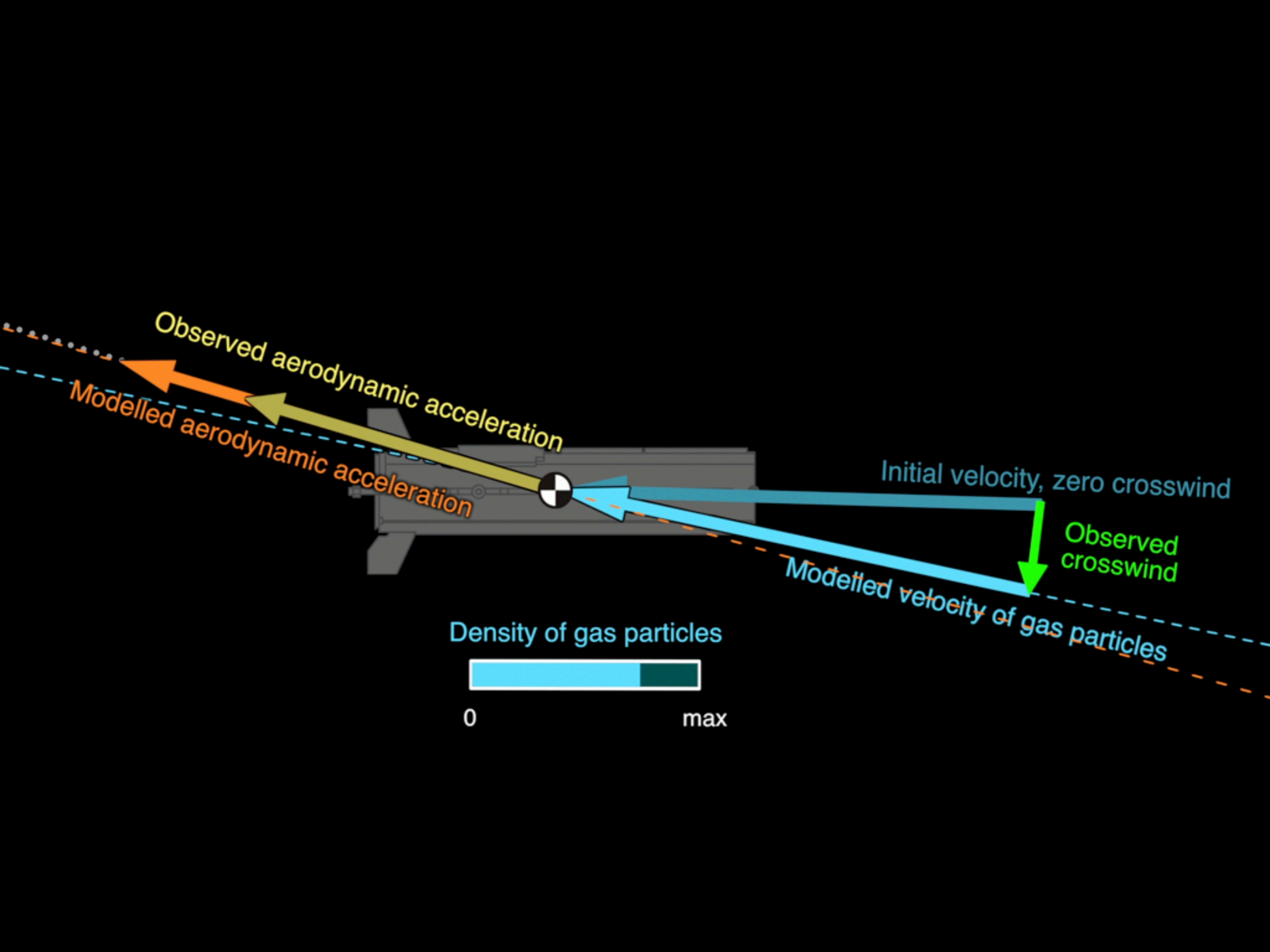


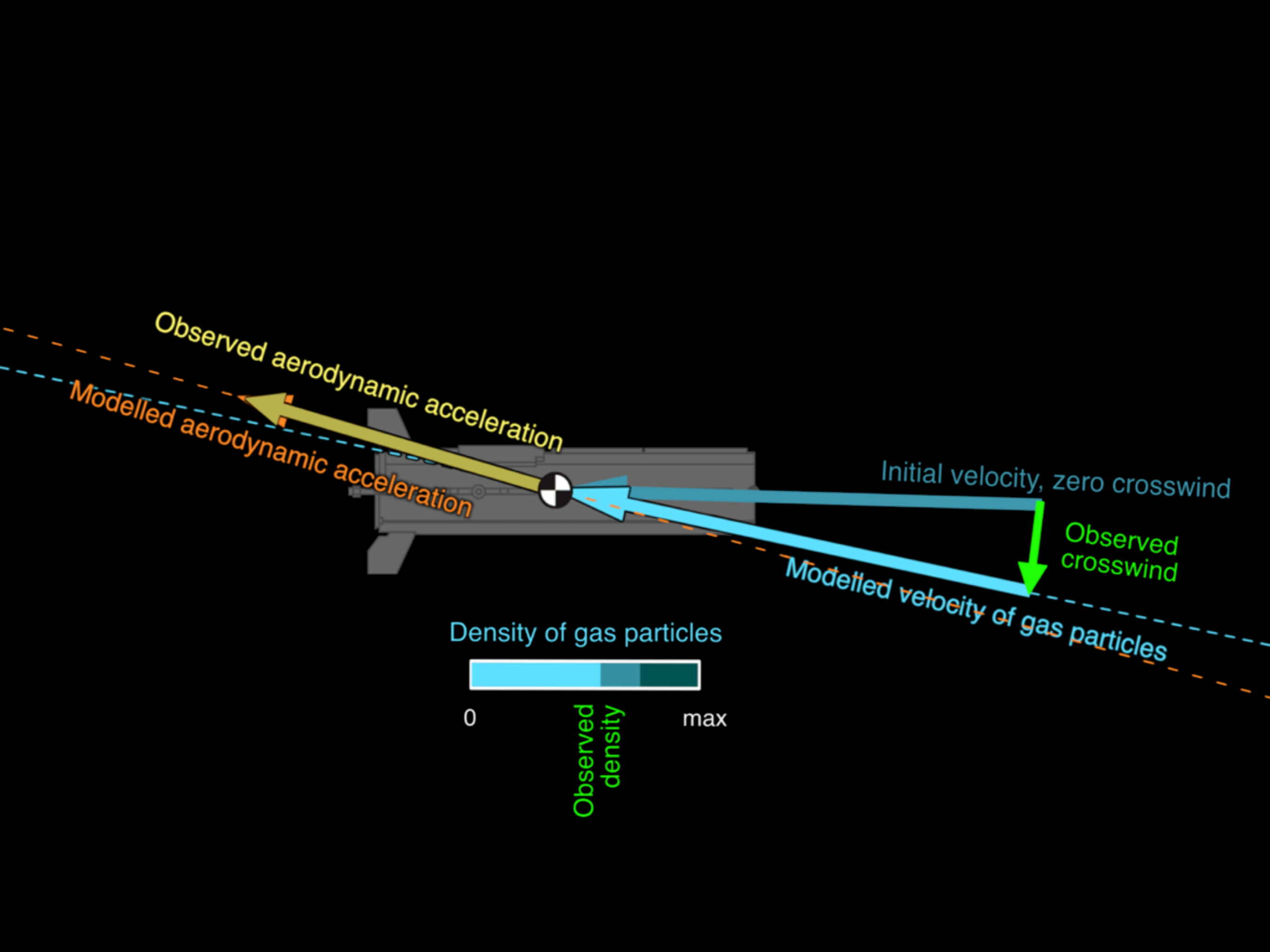
Density of gas particles



0

max





Observed aerodynamic acceleration

Modelled aerodynamic acceleration

Initial velocity, zero crosswind

Observed crosswind

Modelled velocity of gas particles

Density of gas particles

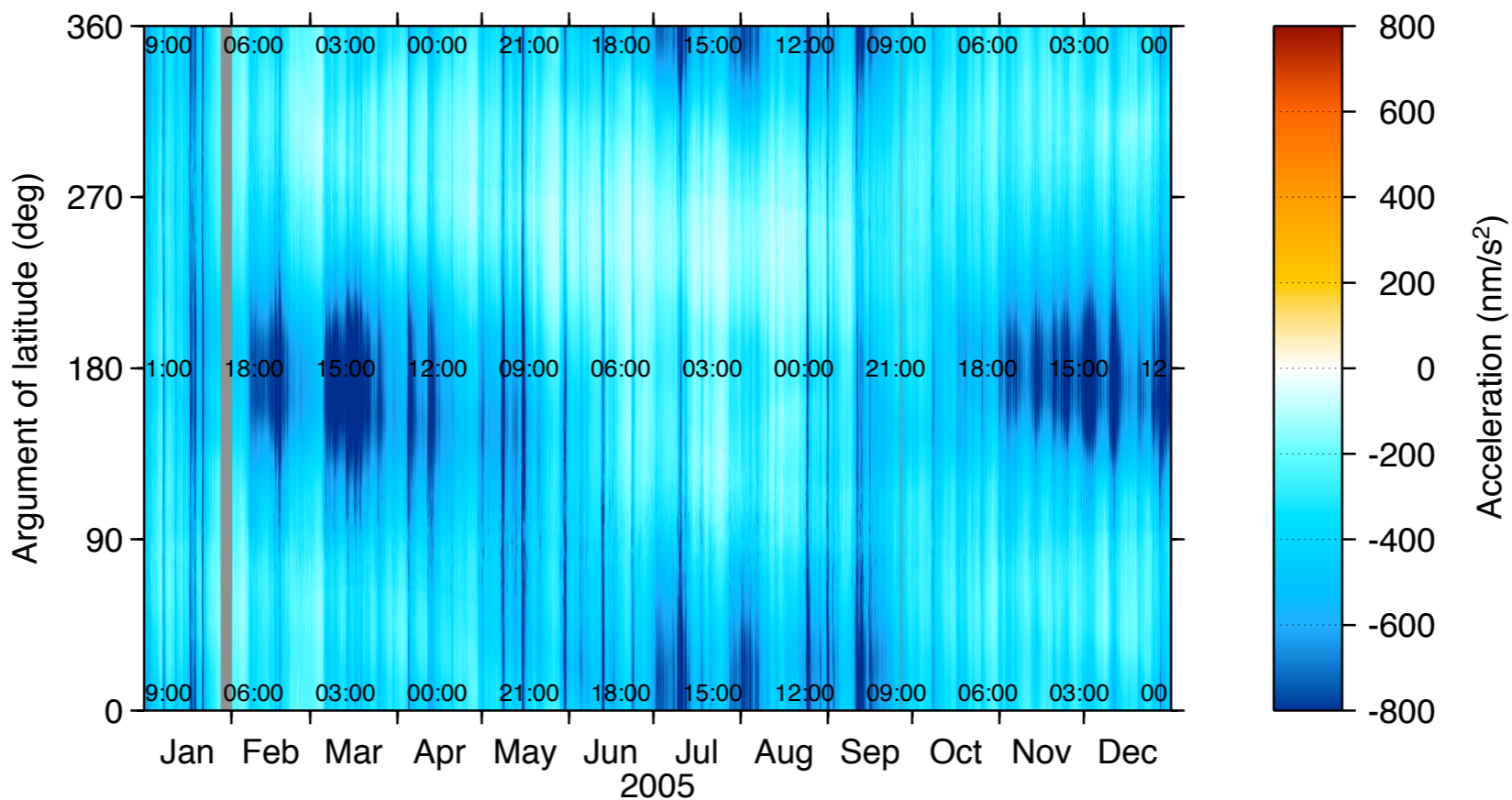


0

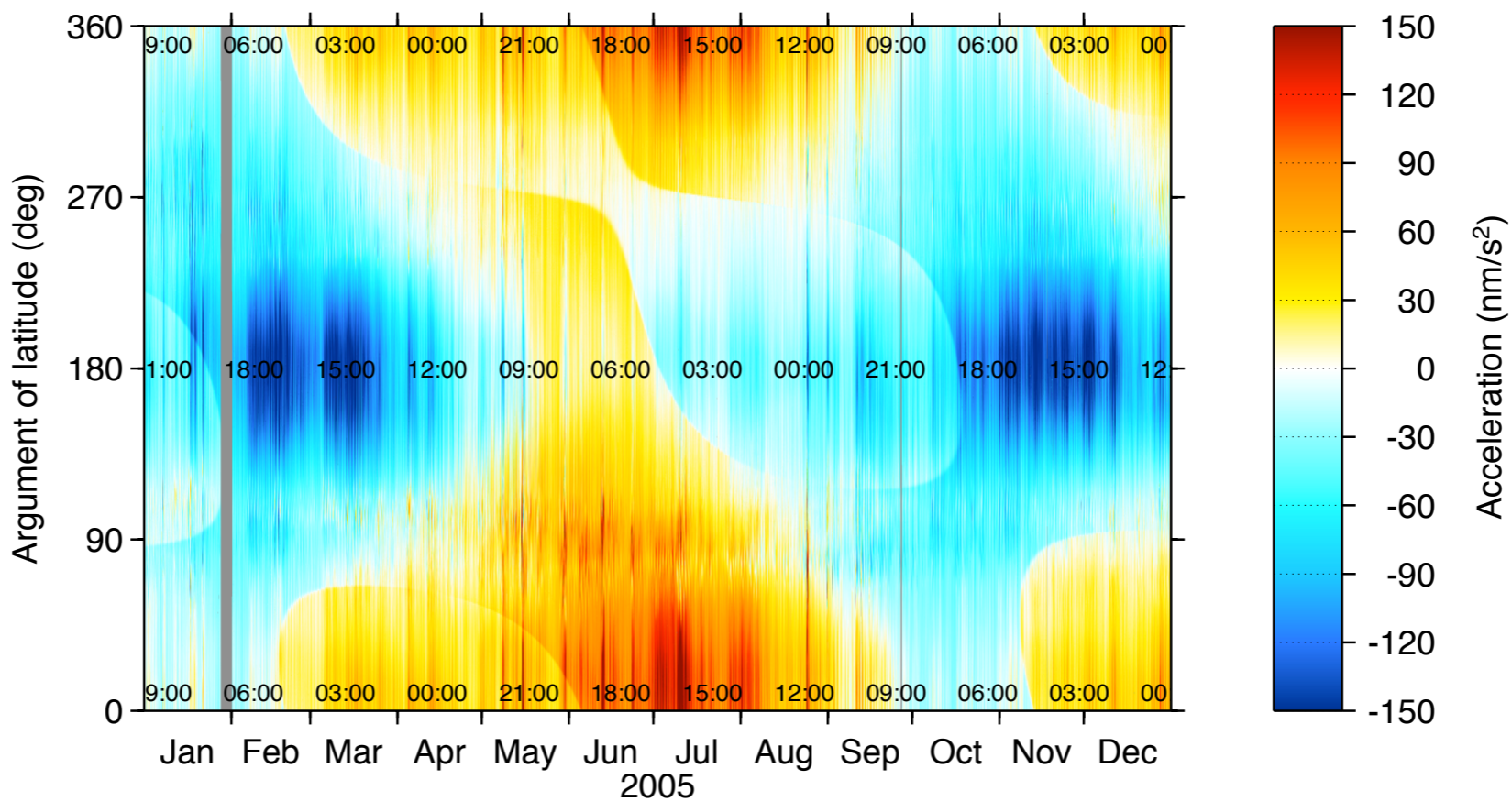
Observed density

max

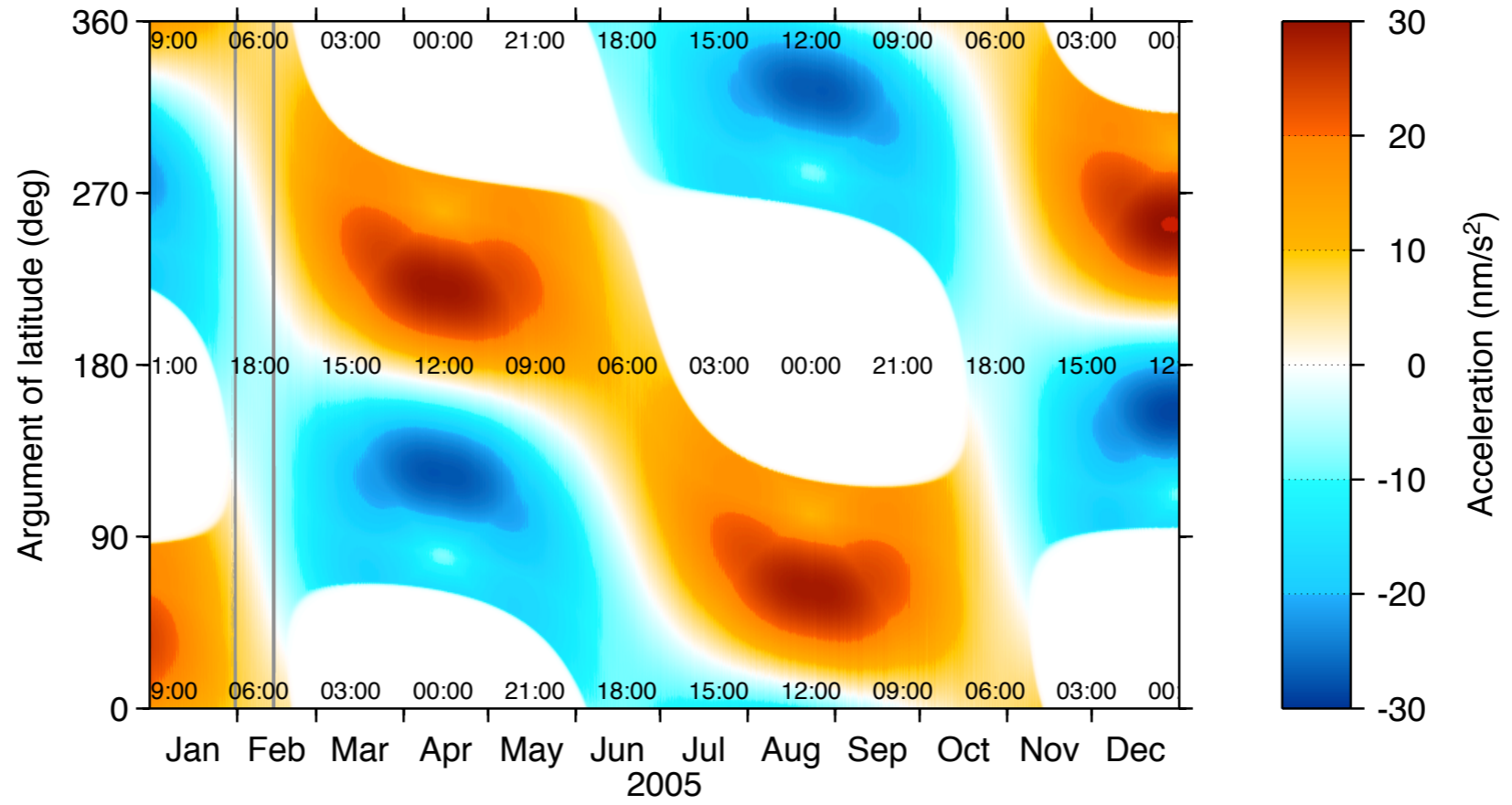
CH_PN_R03B-ACCEL_CALIBRATED_A0.8S-1



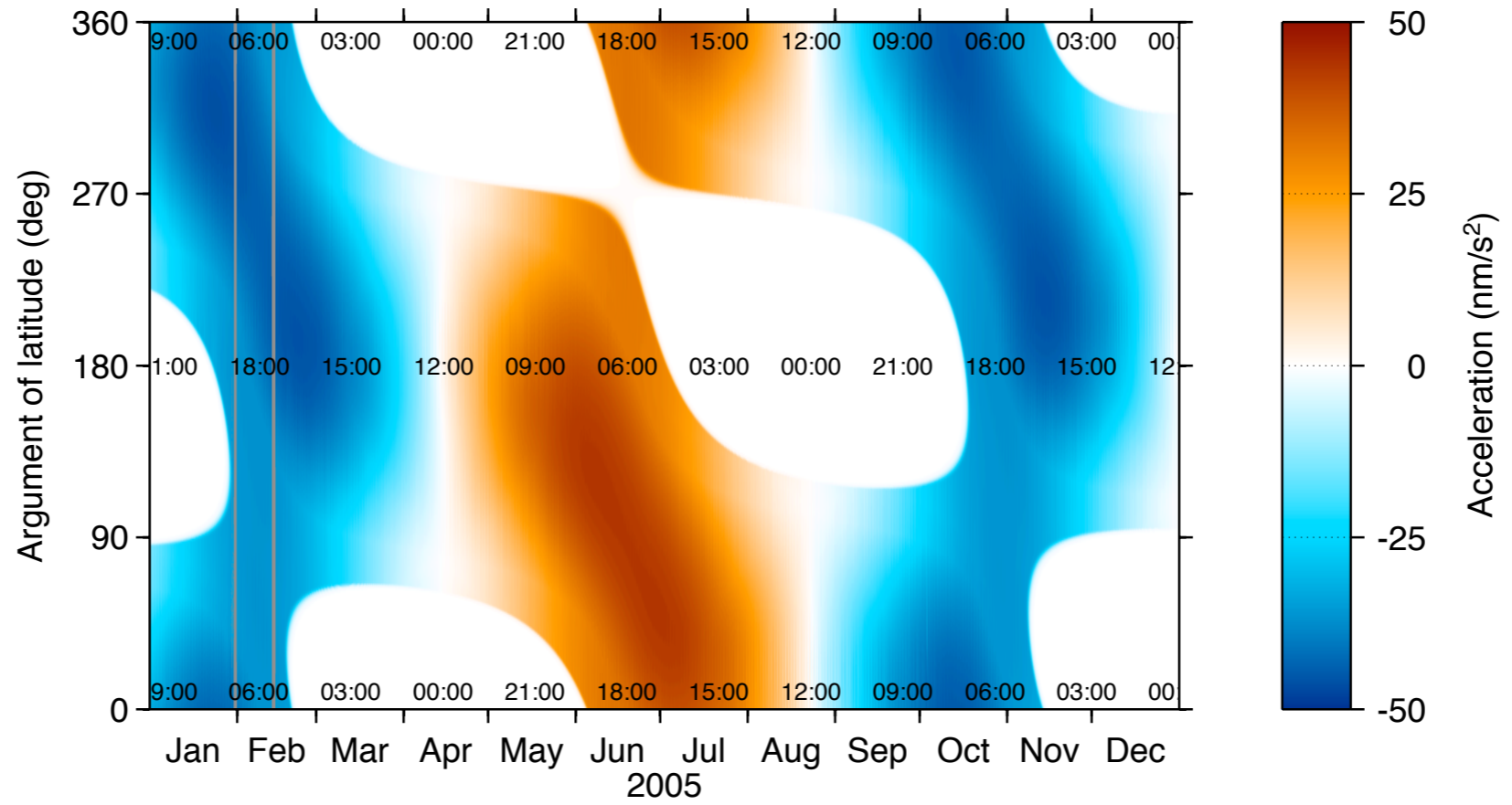
CH_PN_R03B-ACCEL_CALIBRATED_A0.8S-2



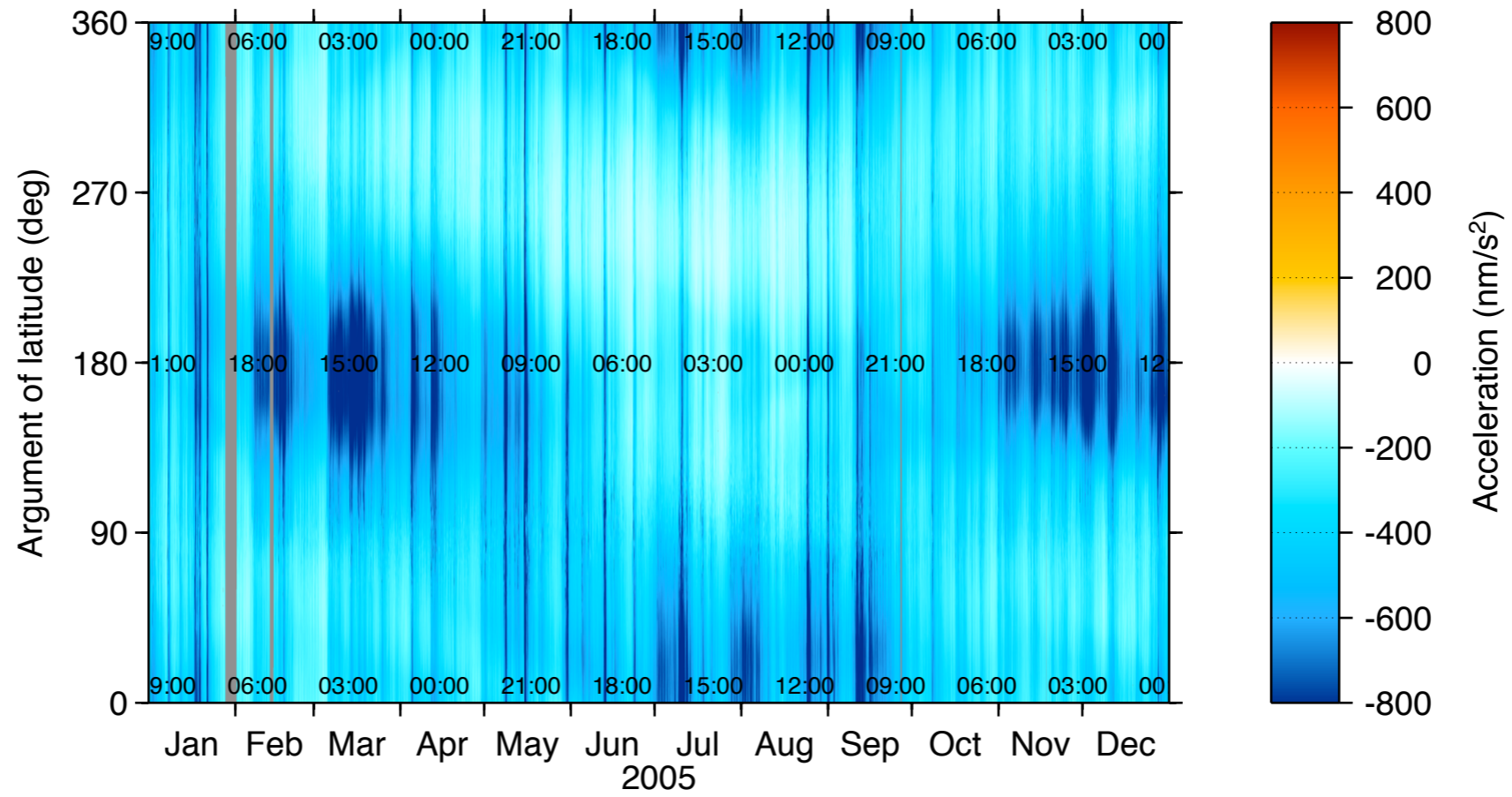
CH_PN-RADPRESS-1



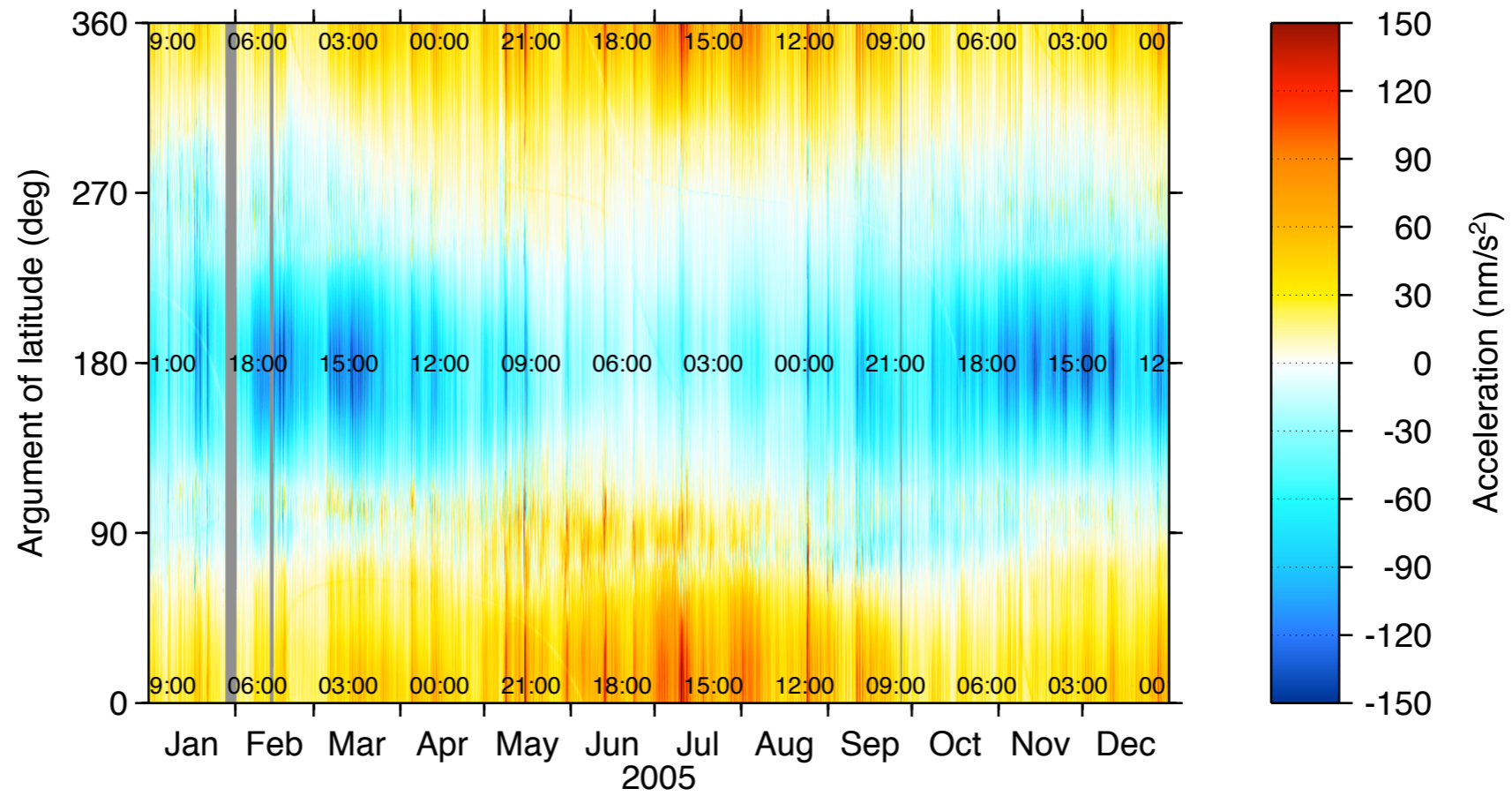
CH_PN-RADPRESS-2

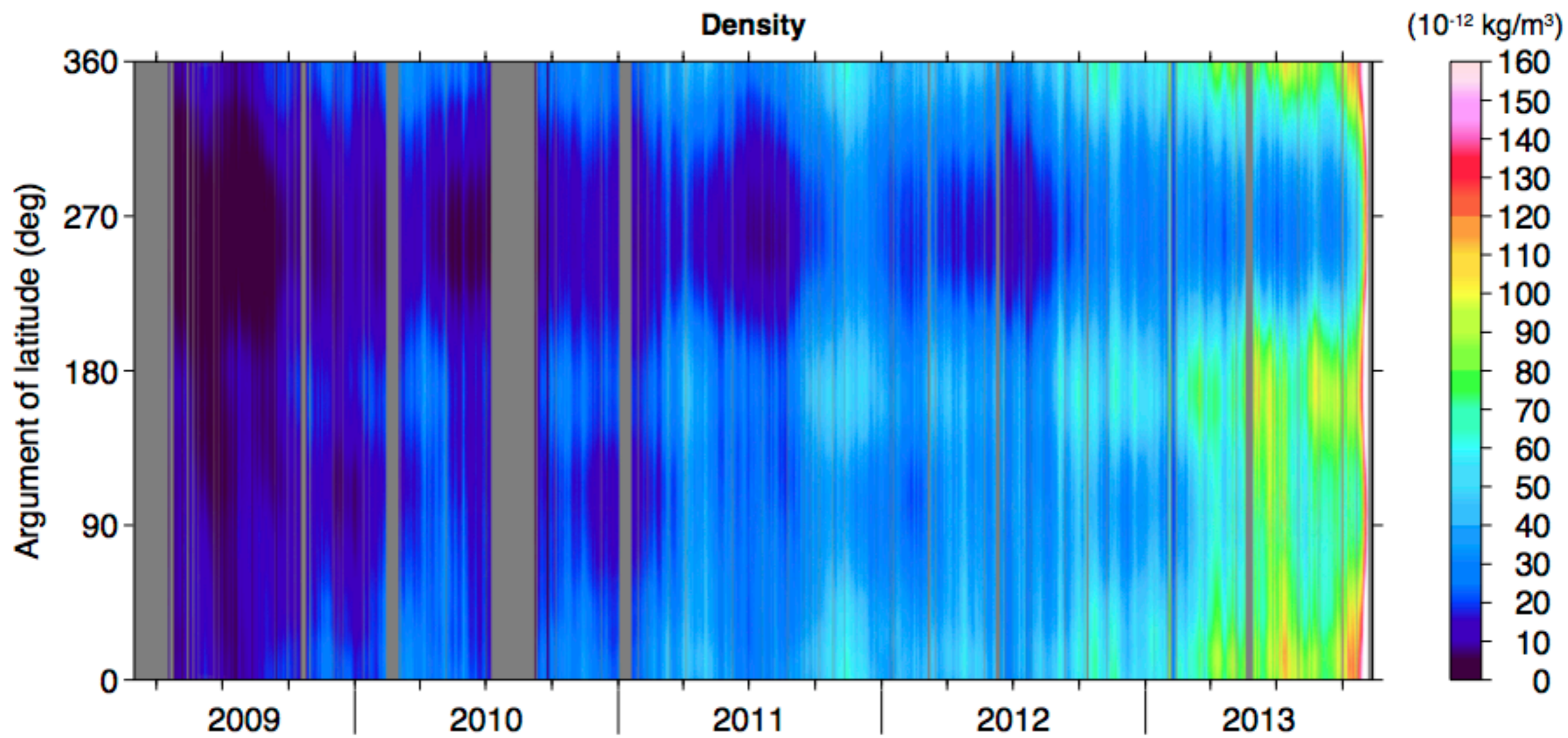


CH_PN_R03B-ACCEL_NORADPRESS_A0.8S-1

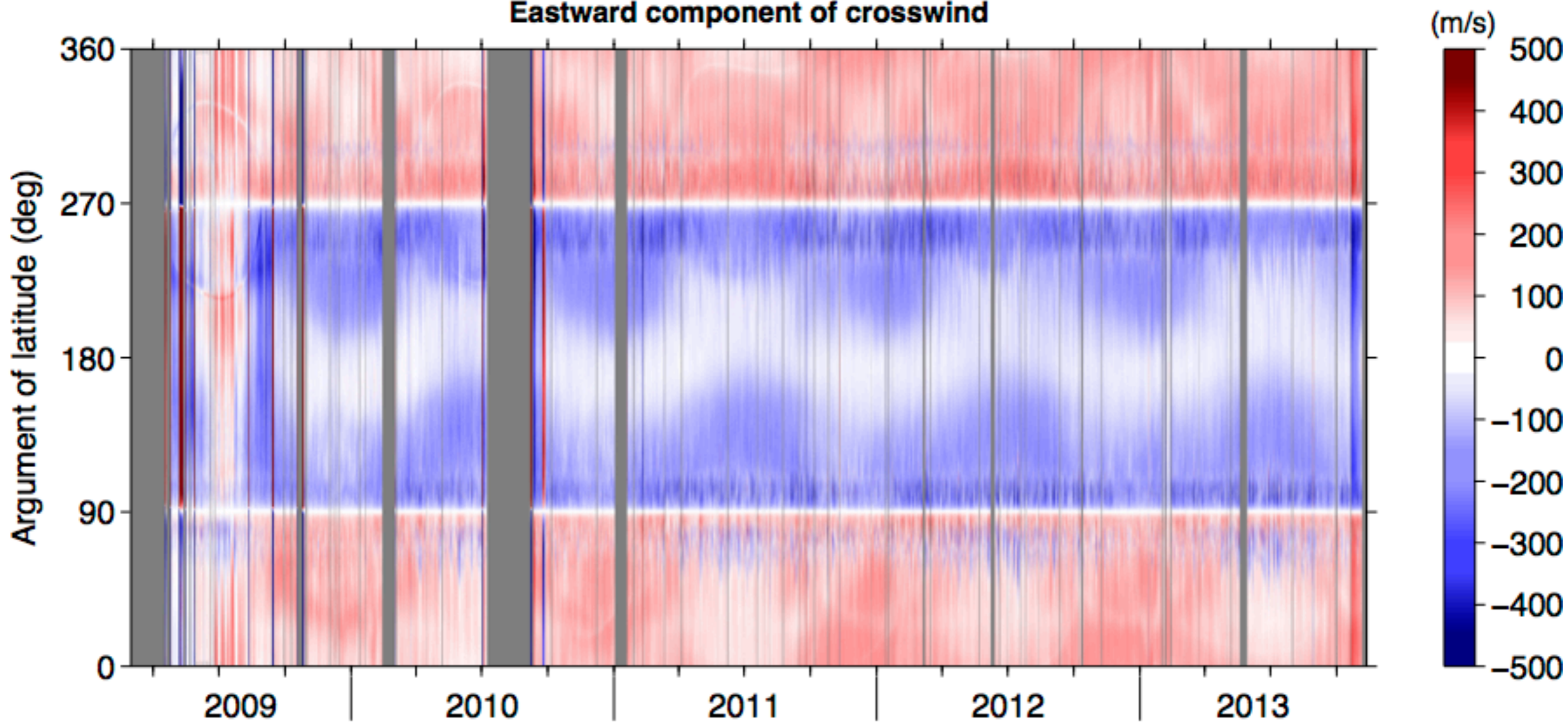


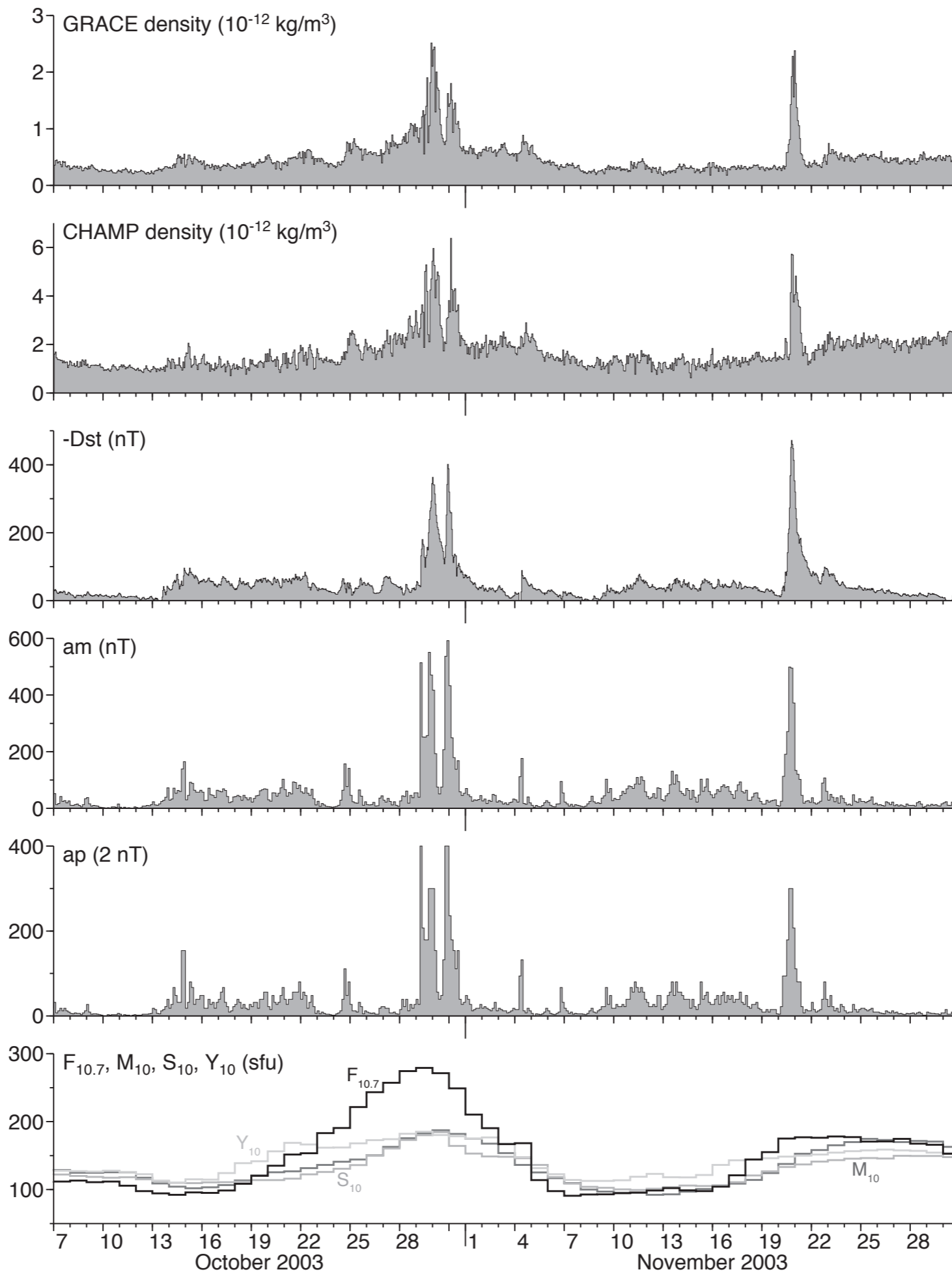
CH_PN_R03B-ACCEL_NORADPRESS_A0.8S-2

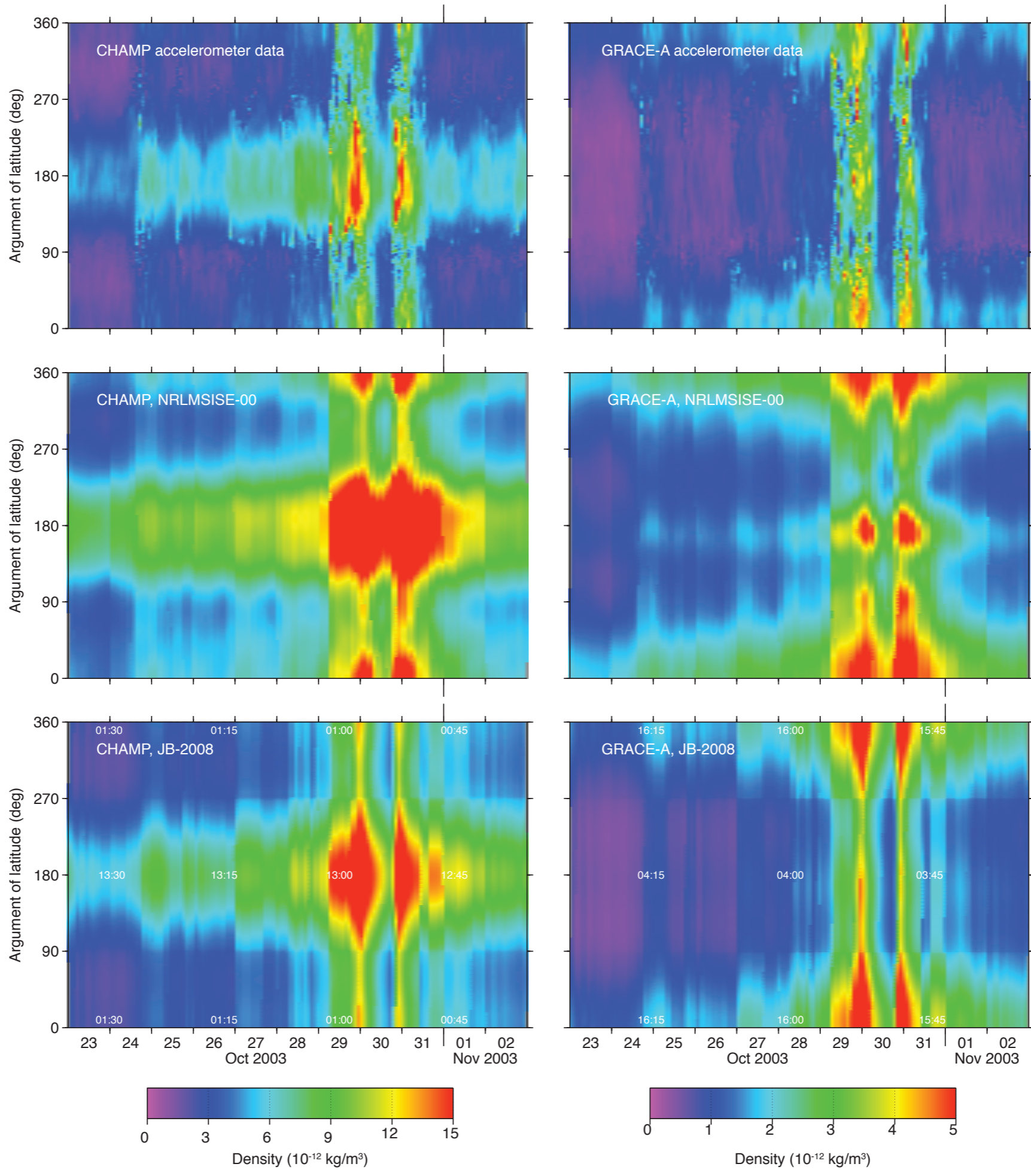




Eastward component of crosswind







Global ionospheric and thermospheric response to the 5 April 2010 geomagnetic storm: An integrated data-model investigation

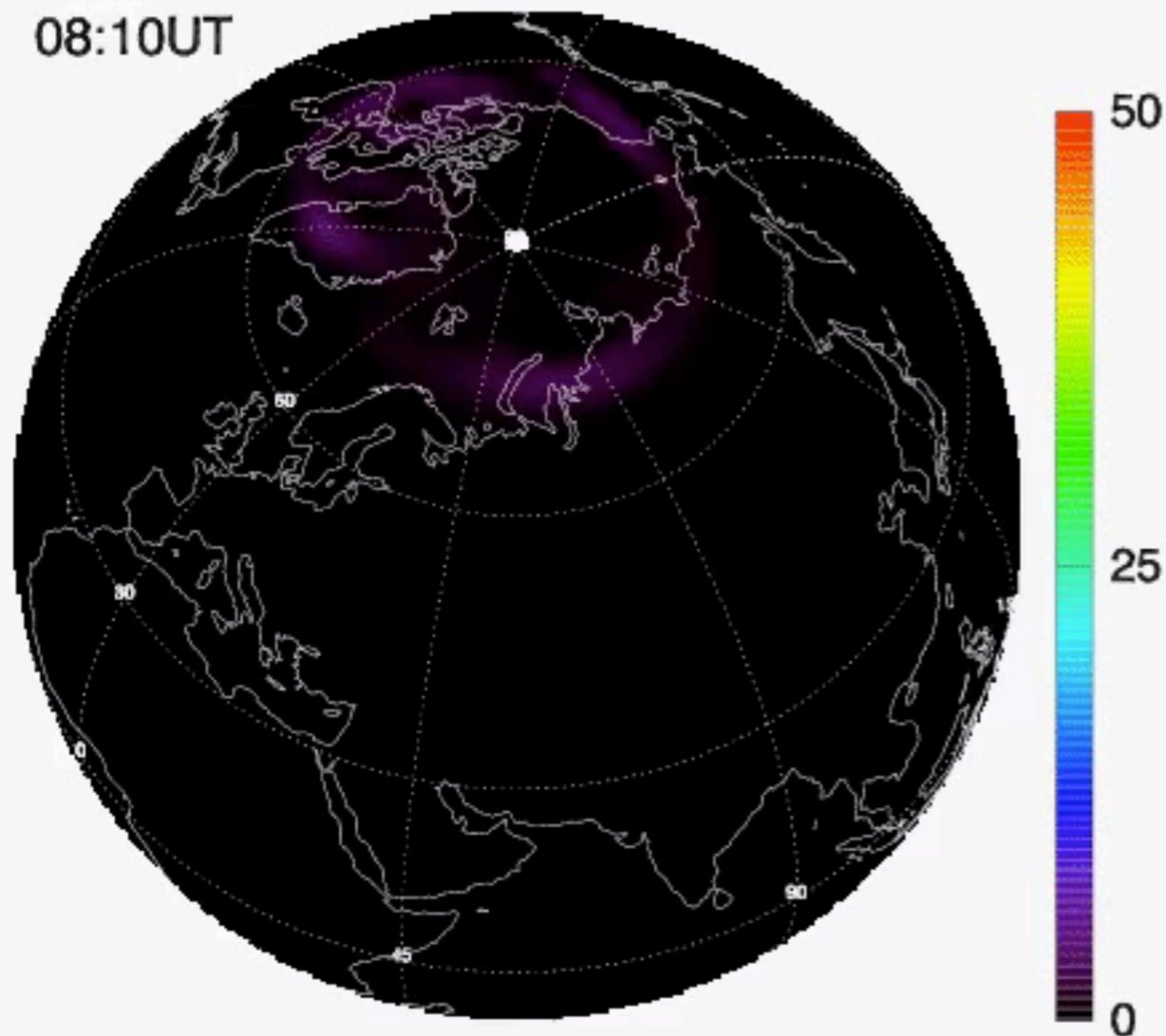
G. Lu¹, M. E. Hagan¹, K. Häusler¹, E. Doornbos², S. Bruinsma³, B. J. Anderson⁴, and H. Korth⁴

Model input from AMPERE, DMSP, SuperDARN & ground magnetometers

Model output

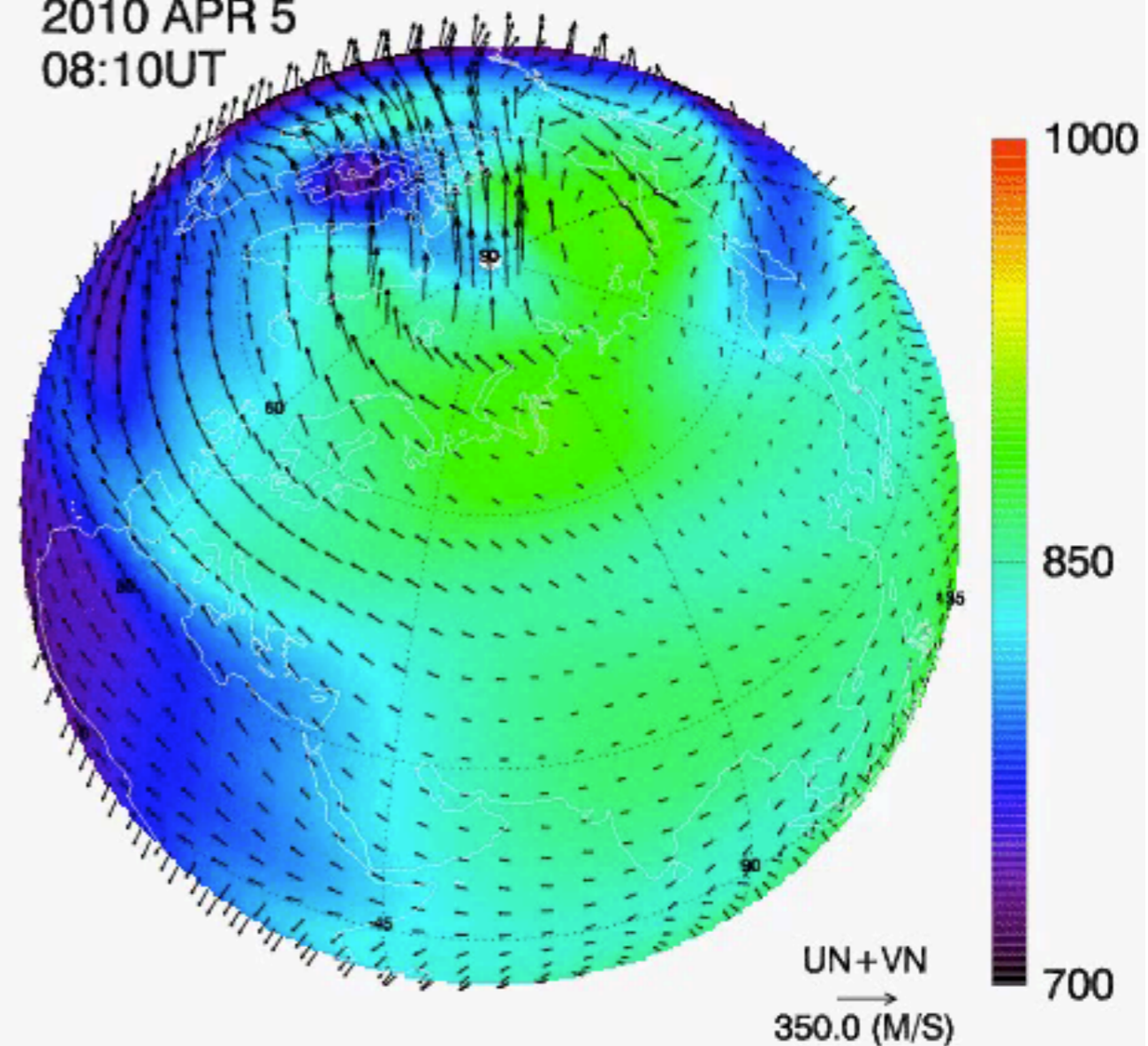
Joule Heat (mW/m^2)

2010 APR 5
08:10UT



Neutral Temperature ($^{\circ}\text{K}$) at 300km

2010 APR 5
08:10UT



Global ionospheric and thermospheric response to the 5 April 2010 geomagnetic storm: An integrated data-model investigation

G. Lu¹, M. E. Hagan¹, K. Häusler¹, E. Doornbos², S. Bruinsma³, B. J. Anderson⁴, and H. Korth⁴

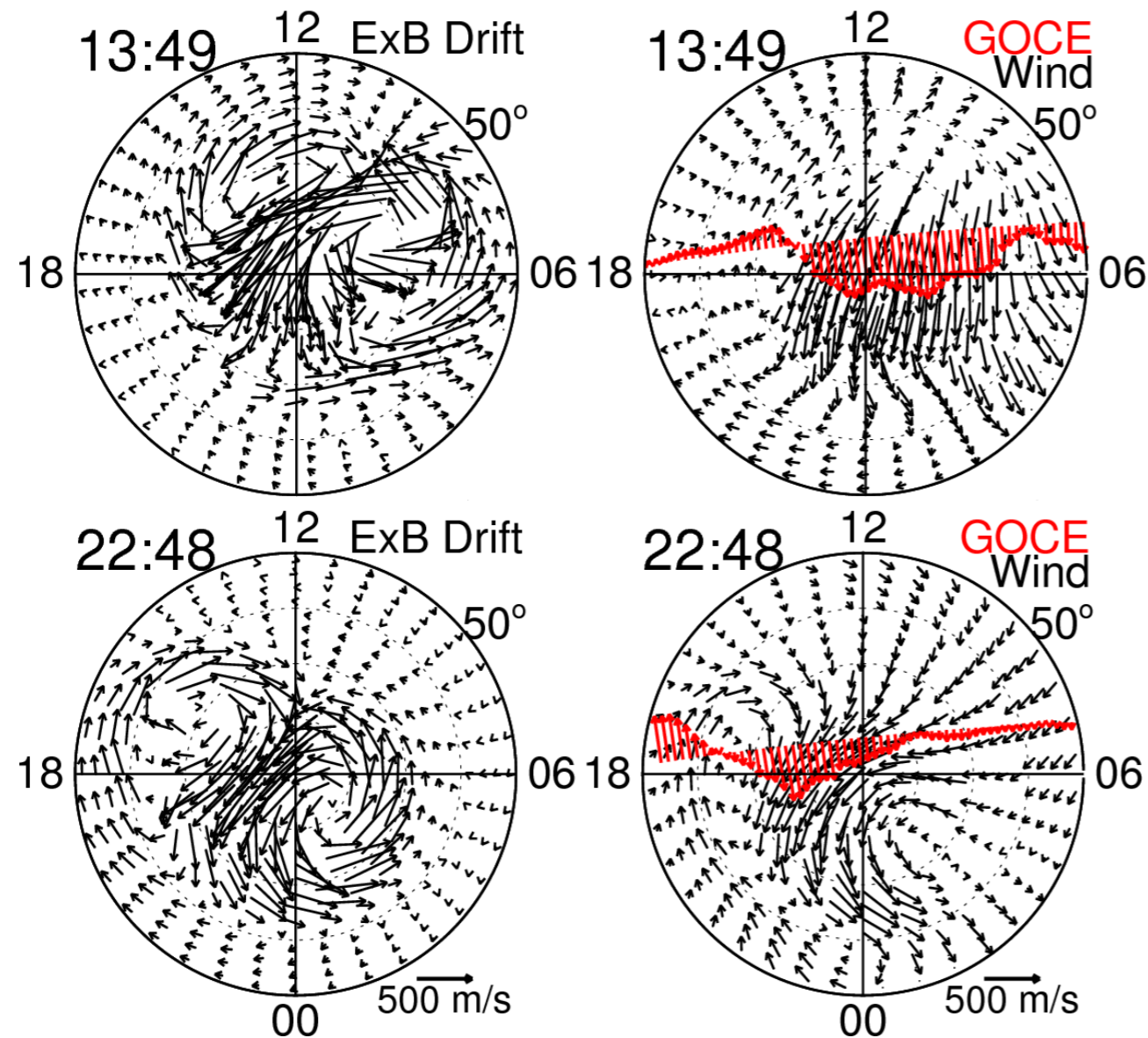
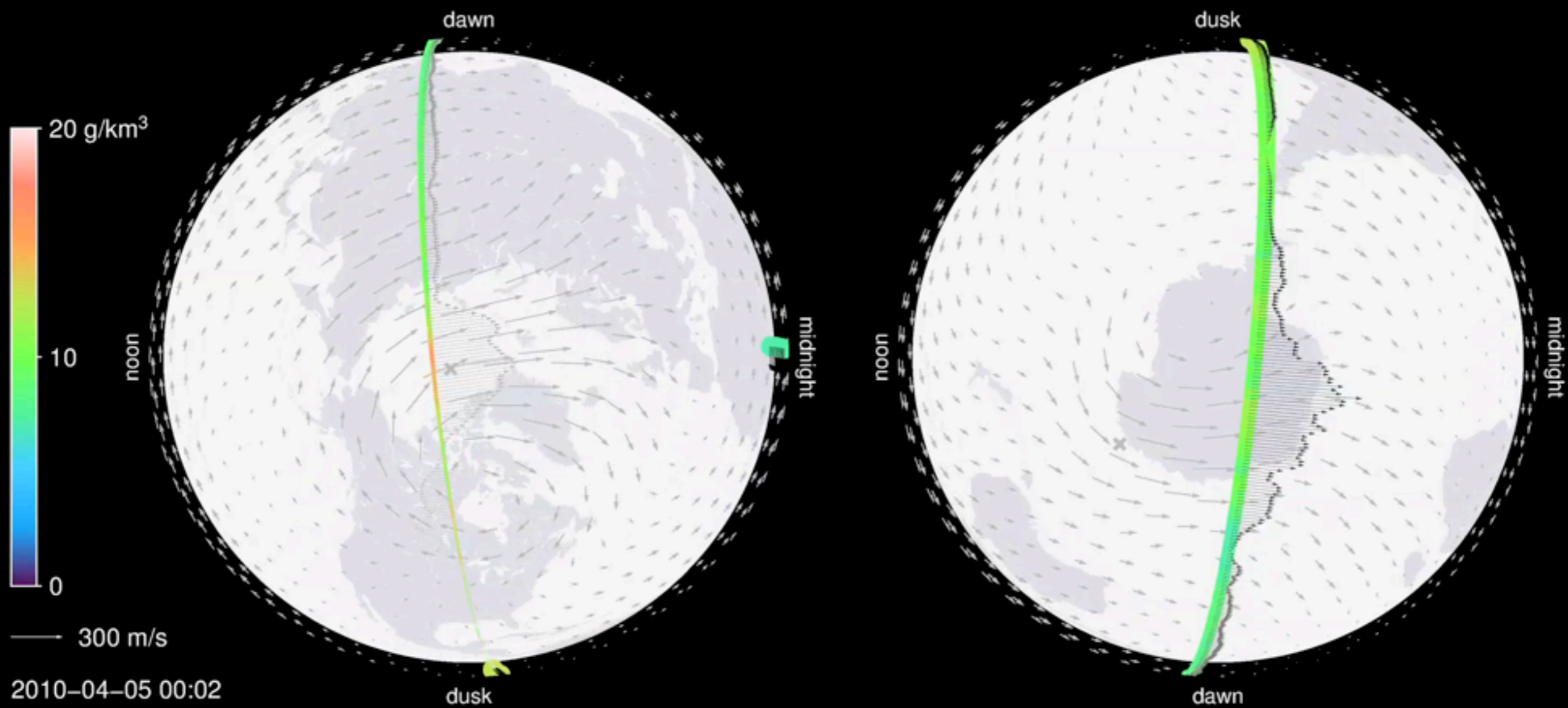


Figure 4. Maps of $\mathbf{E} \times \mathbf{B}$ drifts (left column) and neutral winds (right column) from the TIMEGCM at 250 km in the northern hemisphere plotted in geographic latitude versus local time. The marked UT in each panel corresponds to the time when the GOCE spacecraft was closest to the north pole. The GOCE cross-track winds are plotted in red arrows.

Thermospheric density and wind in the polar regions



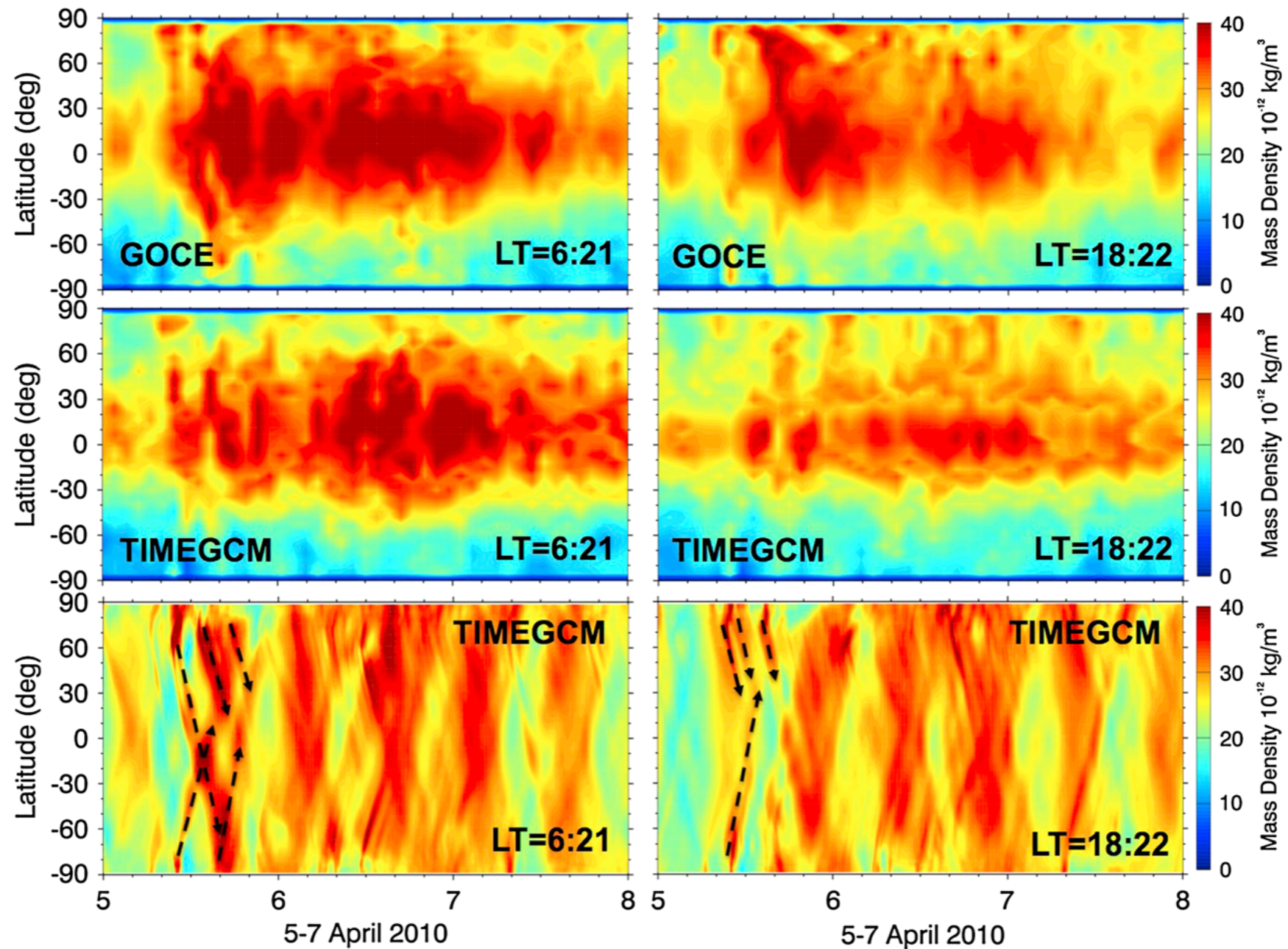


Figure 5. (top row) GOCE neutral mass density measured at (left column) dawn and (right column) dusk. (middle row) Modeled neutral density extracted along the satellite track. (bottom row) Modeled neutral density at fixed local times and a fixed altitude of 266 km. The vertical axis is in geographic latitude. The dashed lines in Figure 5 (bottom row) highlight the main TADs on 5 April with the arrows indicating the TAD propagation directions. The apparent hemispheric difference in Figure 5 (top and middle rows) is due to GOCE's slightly elliptical orbit, which had an average altitude of 266 km in the Northern Hemisphere and 274 km in the Southern Hemisphere.

Air Density Perturbation Measured by GOCE

300 km

+20%

0

-20%

26

27

28

29

30

31

32

33

34

Minutes after earthquake



GOCE satellite
(altitude 270 km)

Speed: 1100 m/s

Wave propagation

Speed: 350 m/s

Thermosphere

85 km

Mesosphere

50 km

Stratosphere

12 km

Troposphere

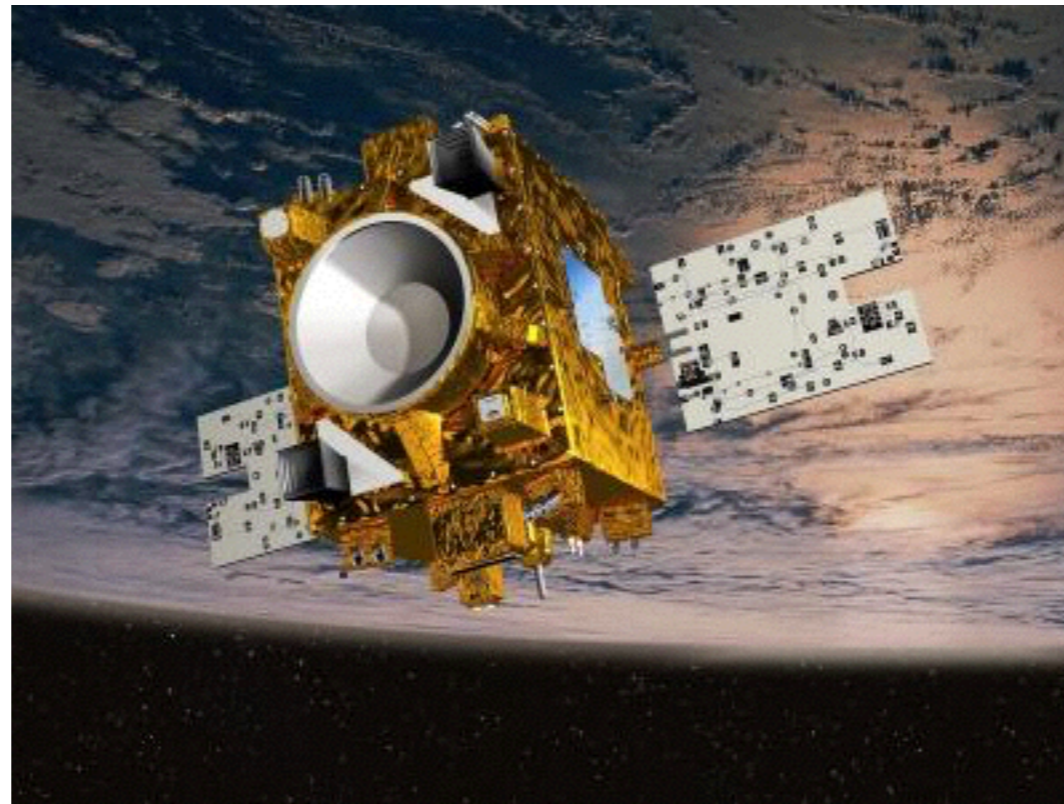
Tōhoku
earthquake
2011



Ongoing activities

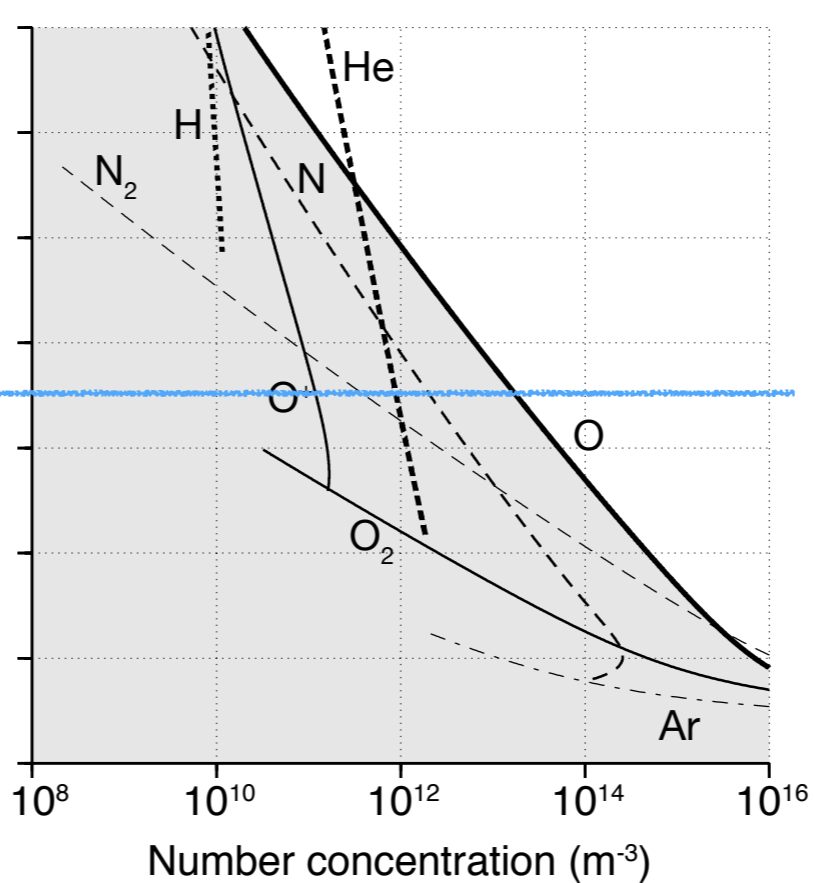
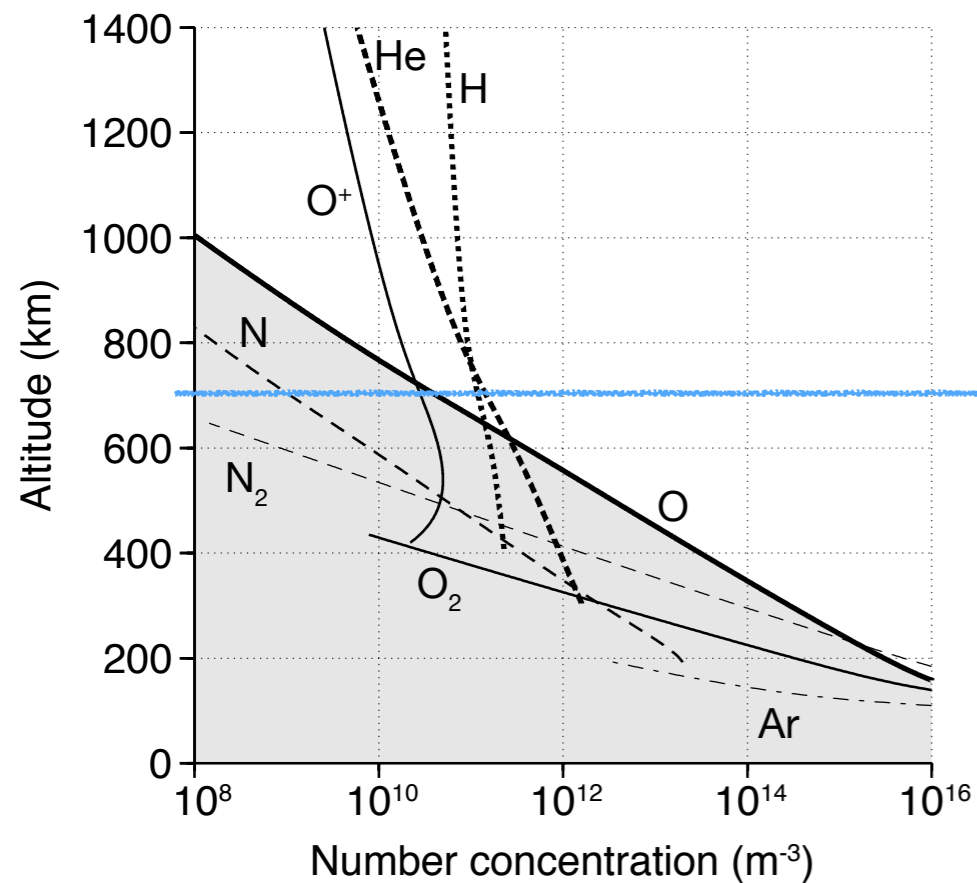
- Characterization of on-board temperature variations on accelerometer signal.
- Elimination of uncertainties in satellite geometry and satellite aerodynamics parameters.
- Use of angular acceleration measurements, analysis of aerodynamic torques using GOCE data.
- Analysis of very low altitude GOCE data at end of life.

Outlook and requirements for MICROSCOPE

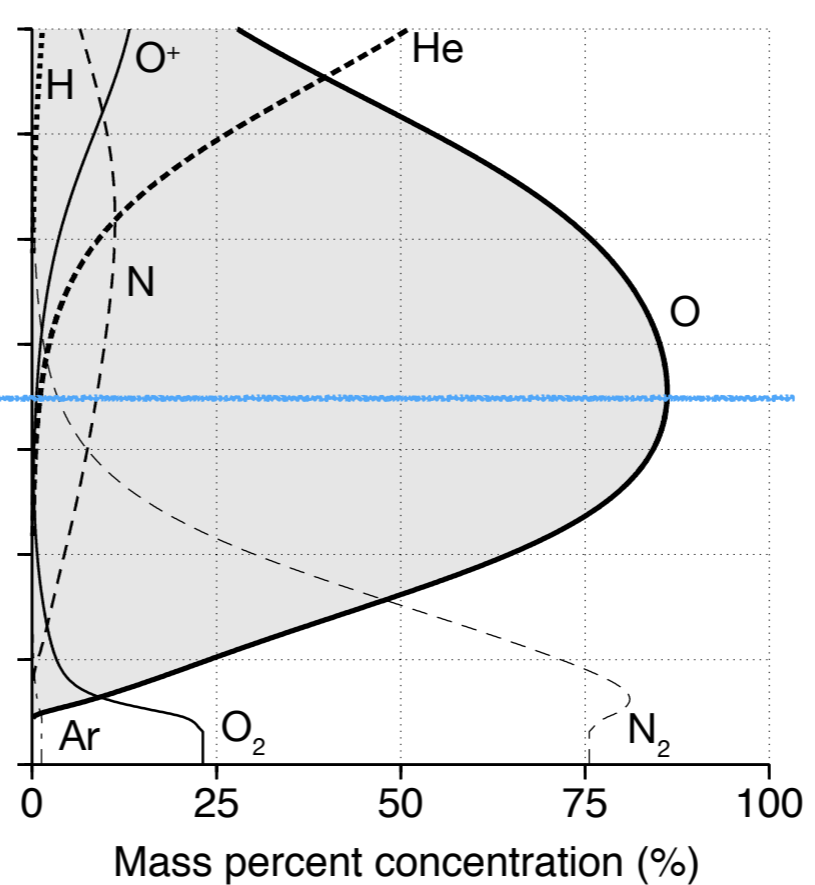
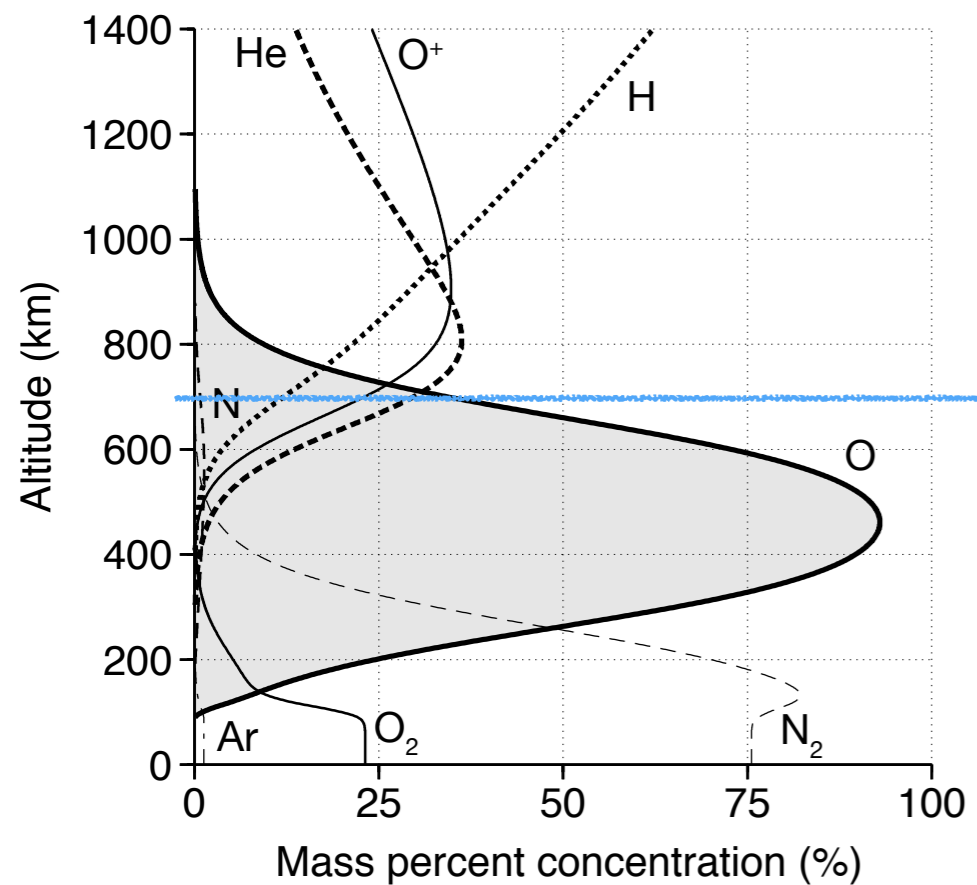


Low activity (2006)

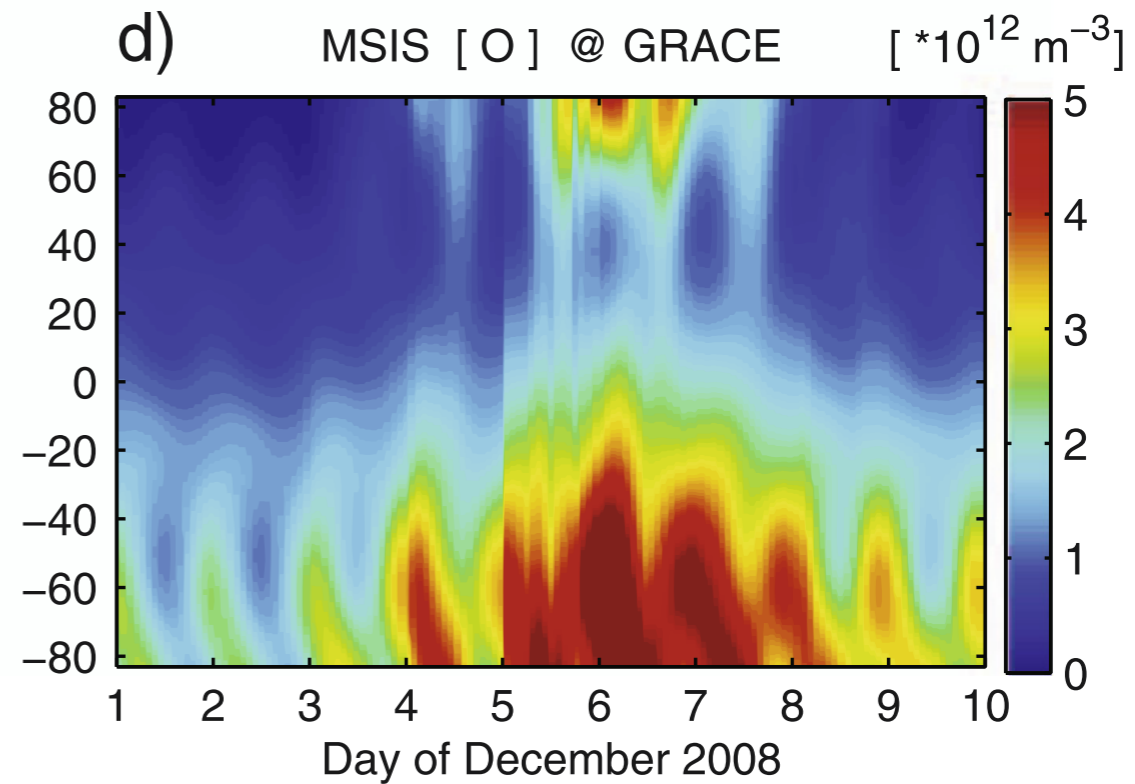
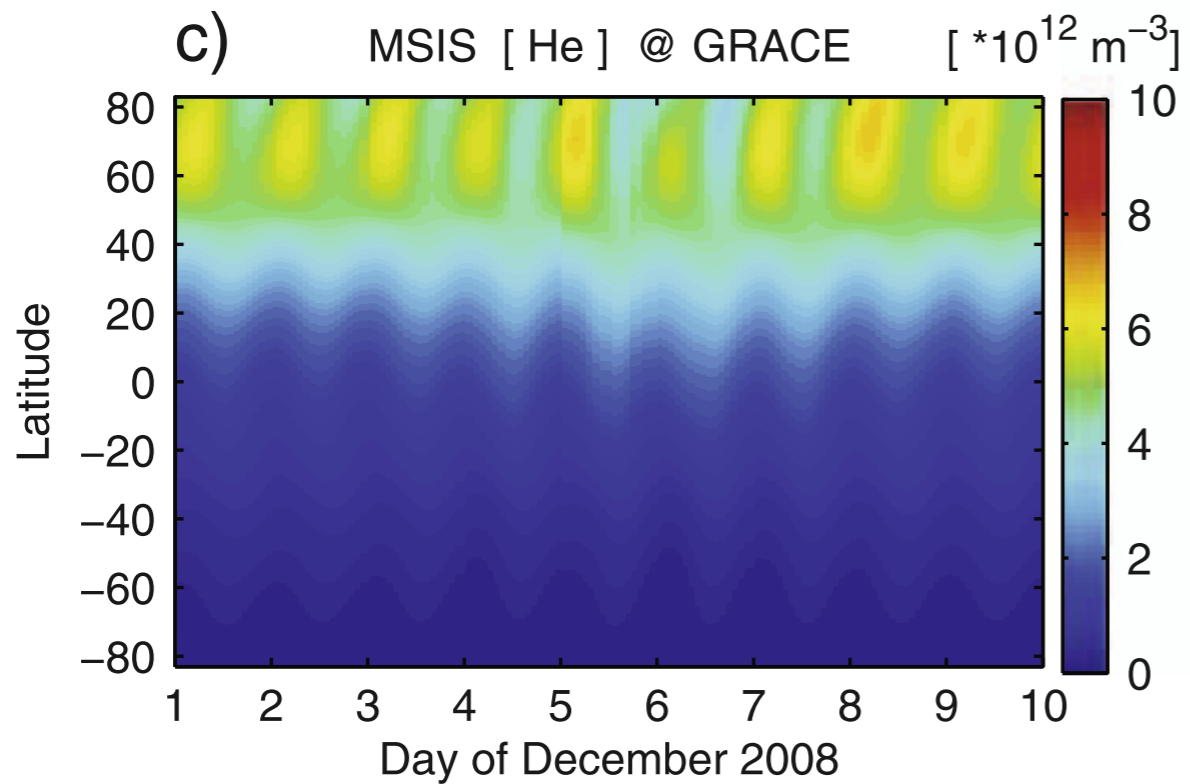
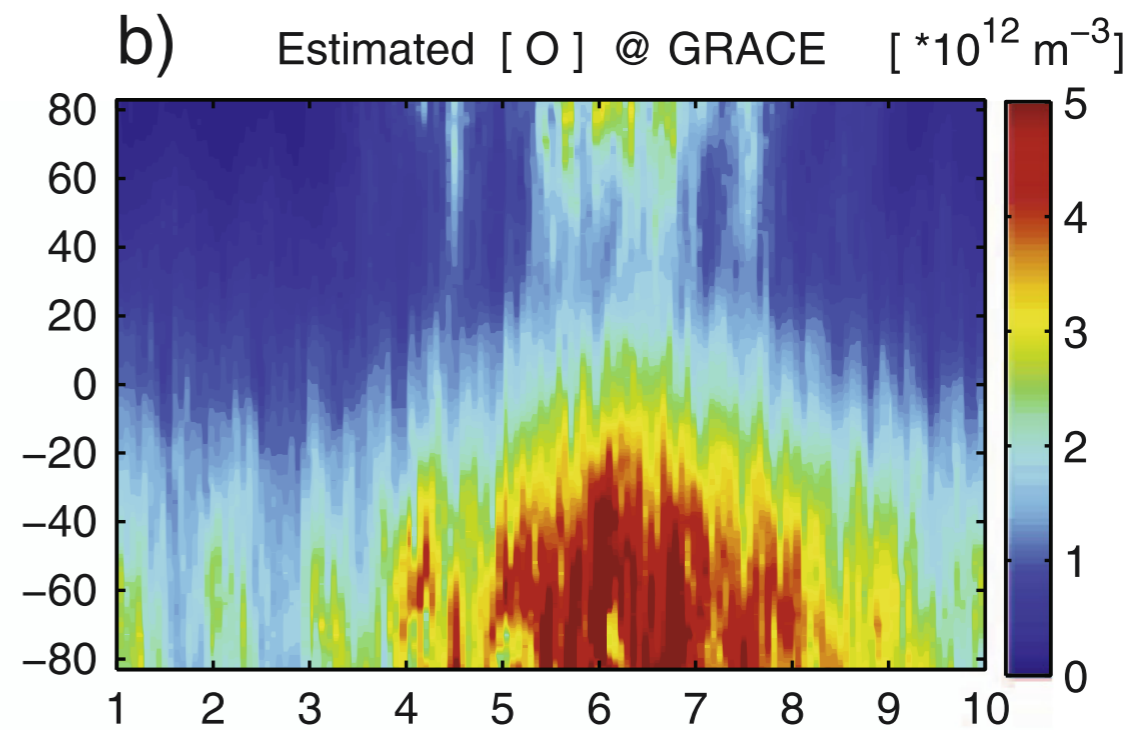
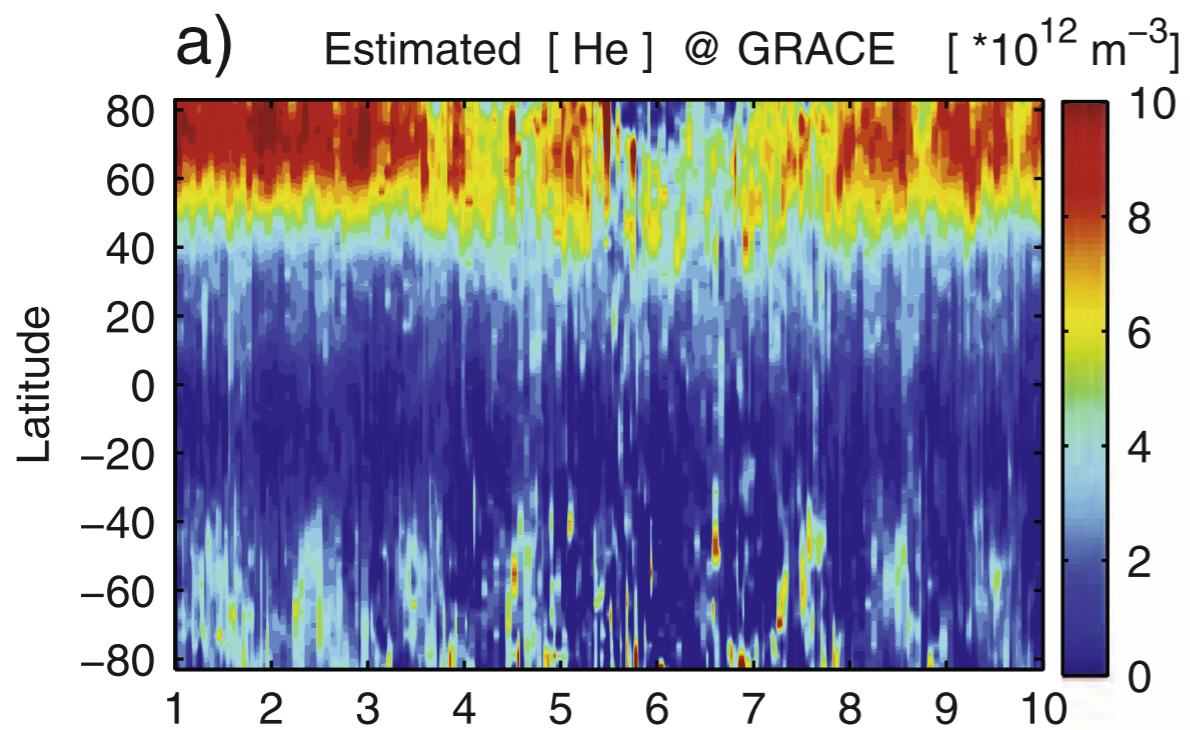
High activity (2000)



MICROSCOPE height



MICROSCOPE height



Characteristics of GOCE and Microscope

GOCE

Operations: 2009-2013

Sun-synchronous 230-270 km

Accelerometers: 10^{-12} m/s²

Drag free in X (flight) direction using ion engine

Microscope

Launch: 2016

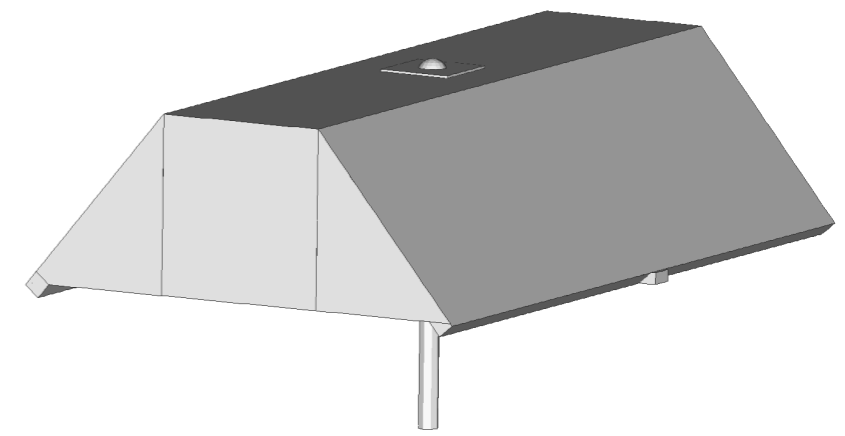
Sun-synchronous 707 km

Accelerometers: 10^{-12} m/s²

Drag free in all directions using micro-thrusters

Requested data

- Accelerations in SI units, if possible uncalibrated & calibrated
 - From accelerometer
 - From AACCS (thrust levels)
- GPS tracking data, e.g. in RINEX format
- For non-gravitational force modelling:
 - Variations of spacecraft mass over time due to fuel consumption
 - Detailed model of geometry of satellite outer surfaces
 - Description of optical properties of satellite outer surfaces



I.5.3.5 GRACE Macro Model: Surface Properties

The surface properties are summarized in the following table. For each surface, the area, the components of its unit normal in the Satellite Frame, the material, as well as its emissivity and absorptivity/reflectivity coefficients are provided.

Panel	Area (m ²)	Unit Normal			Material	Emiss (IR)	Absorp (Vis)	Refl (Vis)		Refl (IR)	
		X	Y	Z				Geom	Diff	Geom	Diff
Front	0.9551567	+1.0	0.0	0.0	SiOx/Kapton	0.62	0.34	0.40	0.26	0.23	0.15
Rear	0.9551567	-1.0	0.0	0.0	SiOx/Kapton	0.62	0.34	0.40	0.26	0.23	0.15
Starboard (outer)	3.1554792	0.0	+0.766044	-0.642787	Si Glass Solar Array	0.81	0.65/0.72 (note 2)	0.05	0.30	0.03	0.16
Starboard (inner)	0.2282913	0.0	-0.766044	+0.642787	SiOx/Kapton	0.62	0.34	0.40	0.26	0.23	0.15
Port (outer)	3.1554792	0.0	-0.766044	-0.642787	Si Glass Solar Array	0.81	0.65/0.72 (note 2)	0.05	0.30	0.03	0.16
Port (inner)	0.2282913	0.0	+0.766044	+0.642787	SiOx/Kapton	0.62	0.34	0.40	0.26	0.23	0.15
Nadir	6.0711120	0.0	0.0	+1.0	Teflon (note 1)	0.75	0.12	0.68	0.20	0.19	0.06
Zenith	2.1673620	0.0	0.0	-1.0	Si Glass Solar Array	0.81	0.65/0.72	0.05	0.30	0.03	0.16
Boom	0.0461901 (note 4)	--	--	--	SiOx/Kapton (note 3)	0.62	0.34	0.40	0.26	0.23	0.15

- (1) fluoro ethylene propylene
- (2) 0.65 for operating solar array (i.e. power being drawn); 0.72 for non-operating array
- (3) S-Band antenna on the boom is protected by a carbon radome (emiss = 0.85; absorp = 0.95), neglected here.
- (4) Planar projection area of the cylindrical Boom, along any direction in the Satellite Frame (X-Y) plane.

Challenges of using MICROSCOPE data for thermosphere studies

- Relatively low aerodynamic acceleration signal.
 - Need to very accurately model radiation pressure.
- Attitude not necessarily kept aligned with flight direction.
 - Need to have good calibration for all accelerometer axes.
- Continuous microthruster operation.
 - Need to characterize the

Benefits of Microscope for thermosphere studies

- Compact satellite shape, easier to model satellite aerodynamics.
- Data at altitude close to the most critical altitude range for spacecraft conjunction assessment and collision avoidance
- High-resolution data on the winter Helium bulge, and its response to the various drivers of the thermosphere.